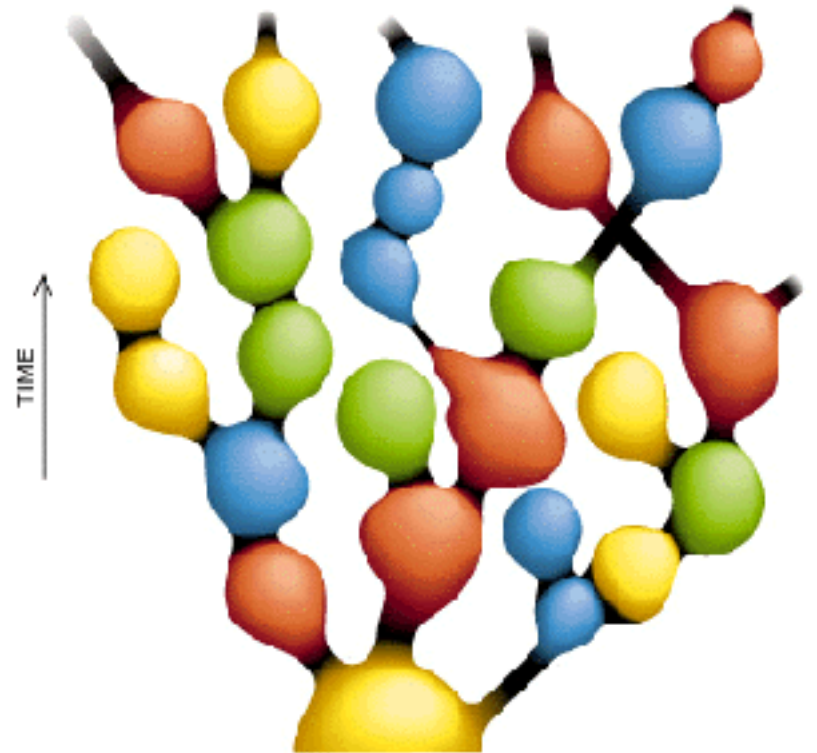
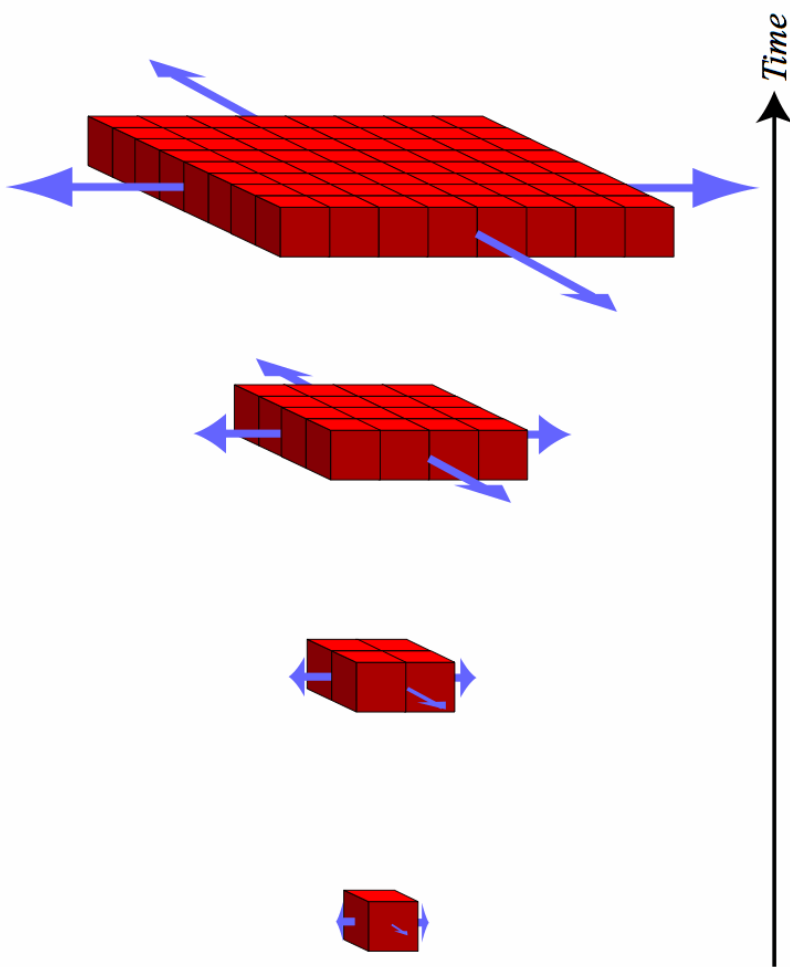
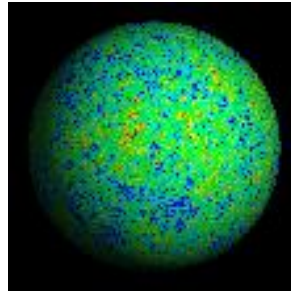


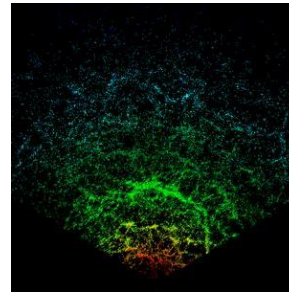
The Inflating Universe



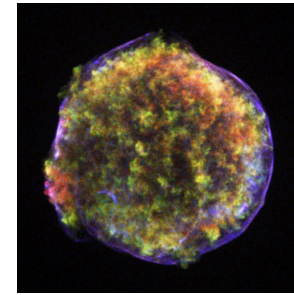
Max Tegmark, MIT



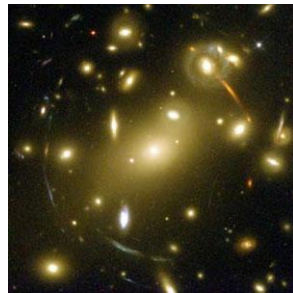
Microwave background



Galaxy surveys

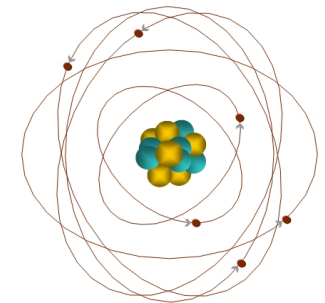


Supernovae Ia

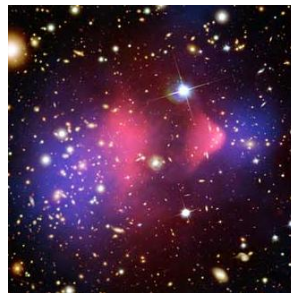


Gravitational lensing

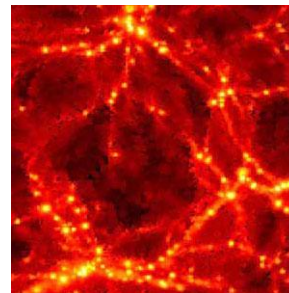
THE COSMIC SMÖRGÅSBORD



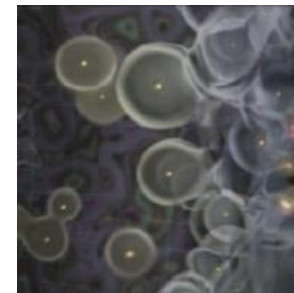
Big Bang nucleosynthesis



Galaxy clusters



Lyman α forest



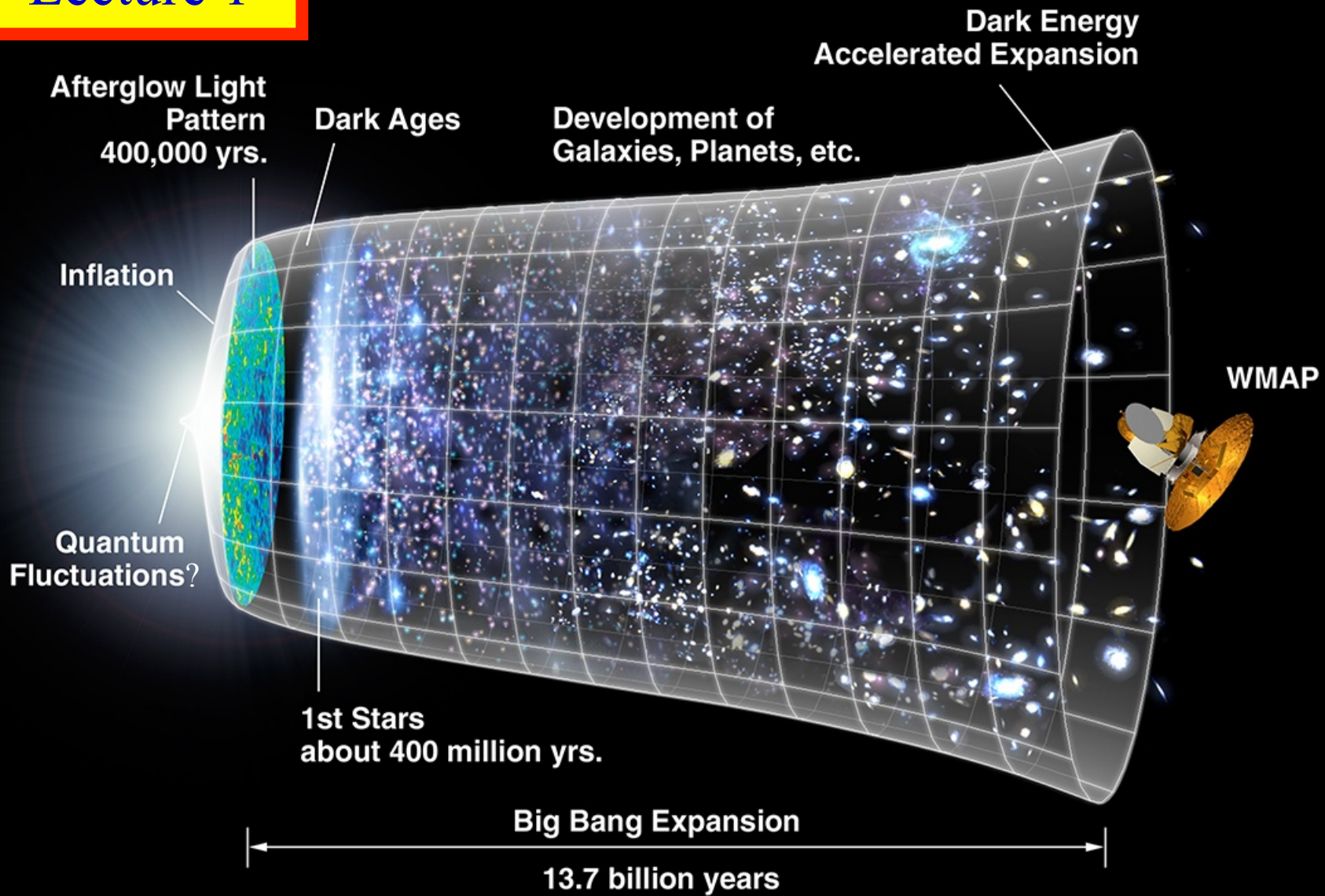
Neutral hydrogen tomography



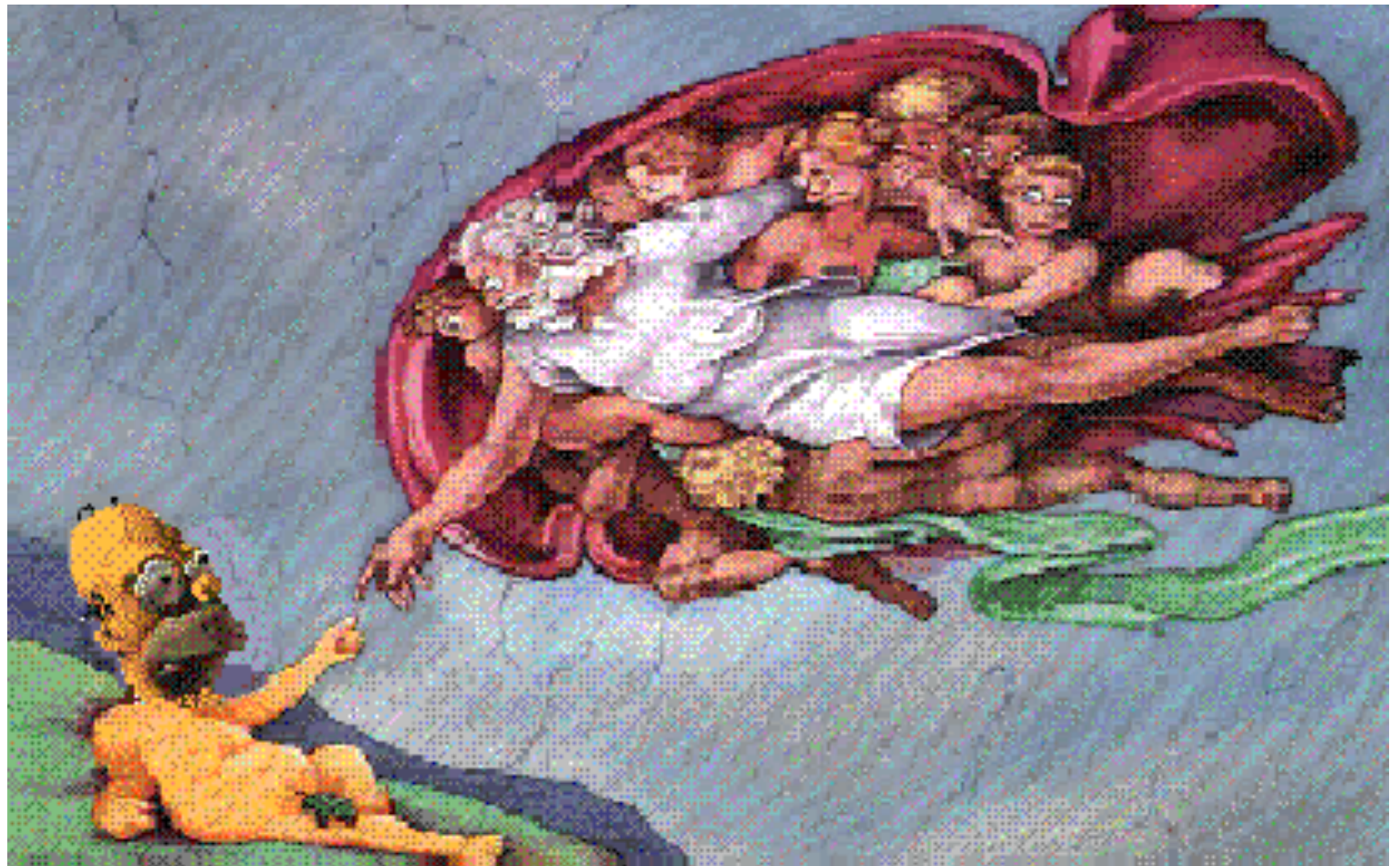
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What have
we learned?

Summary of Lecture 1

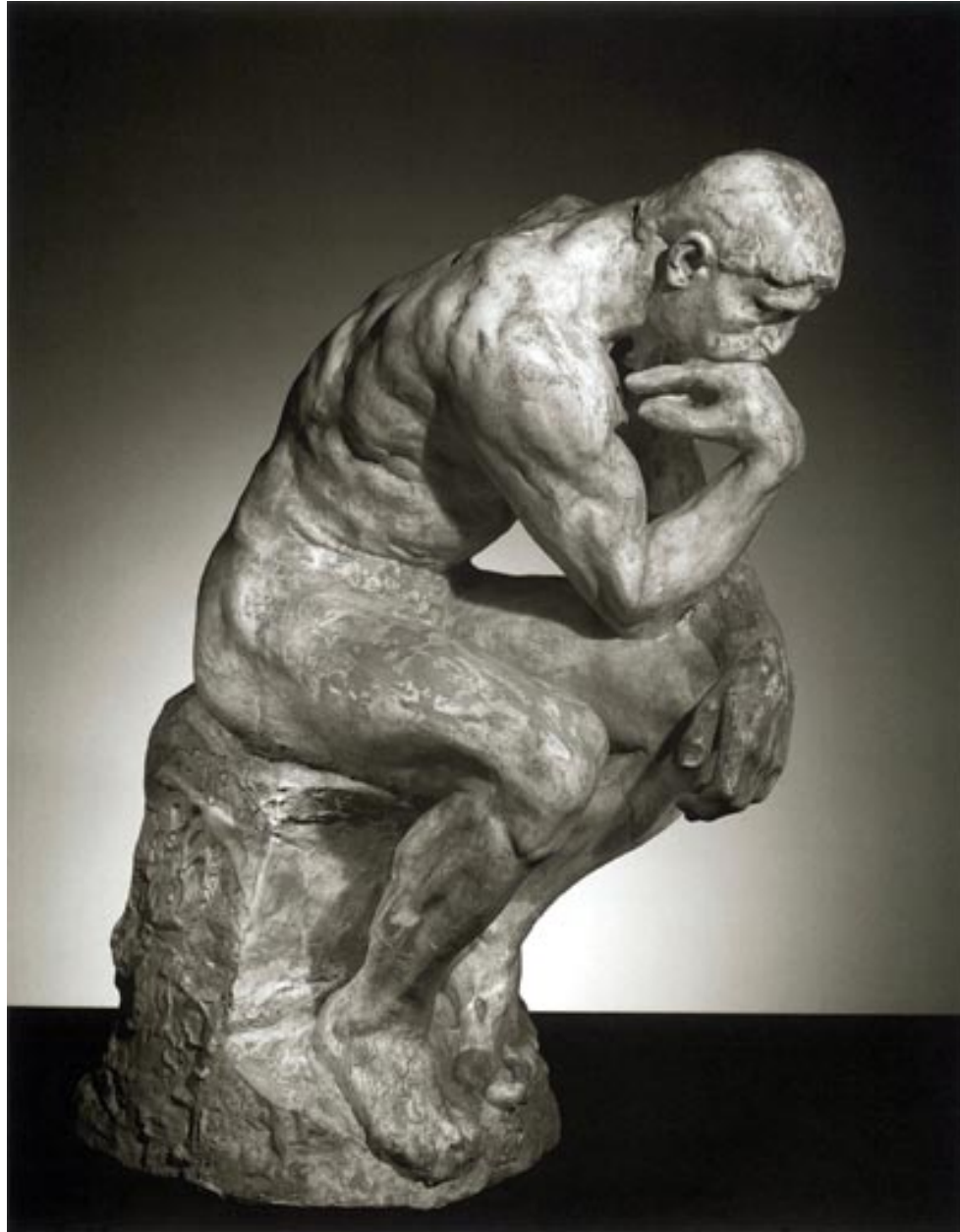


How did it all begin?



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Some things to ponder...

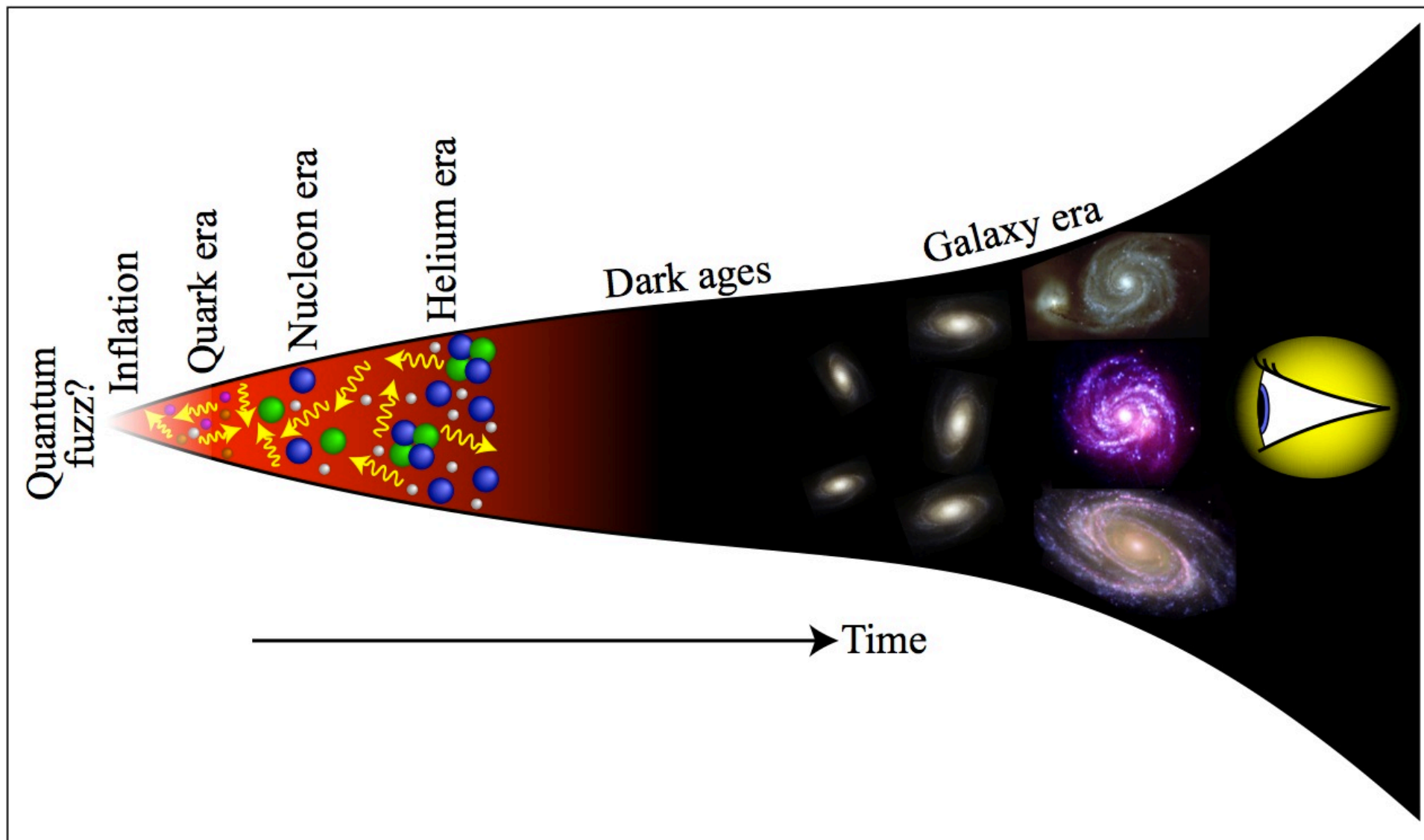


Alan
Guth

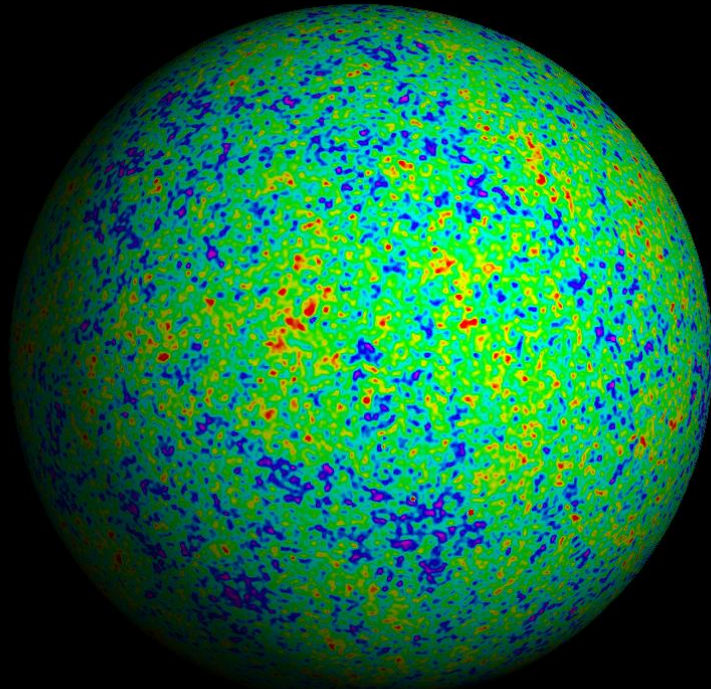


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What put the “bang” into our Big Bang?

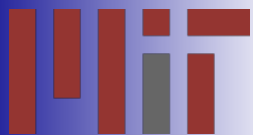
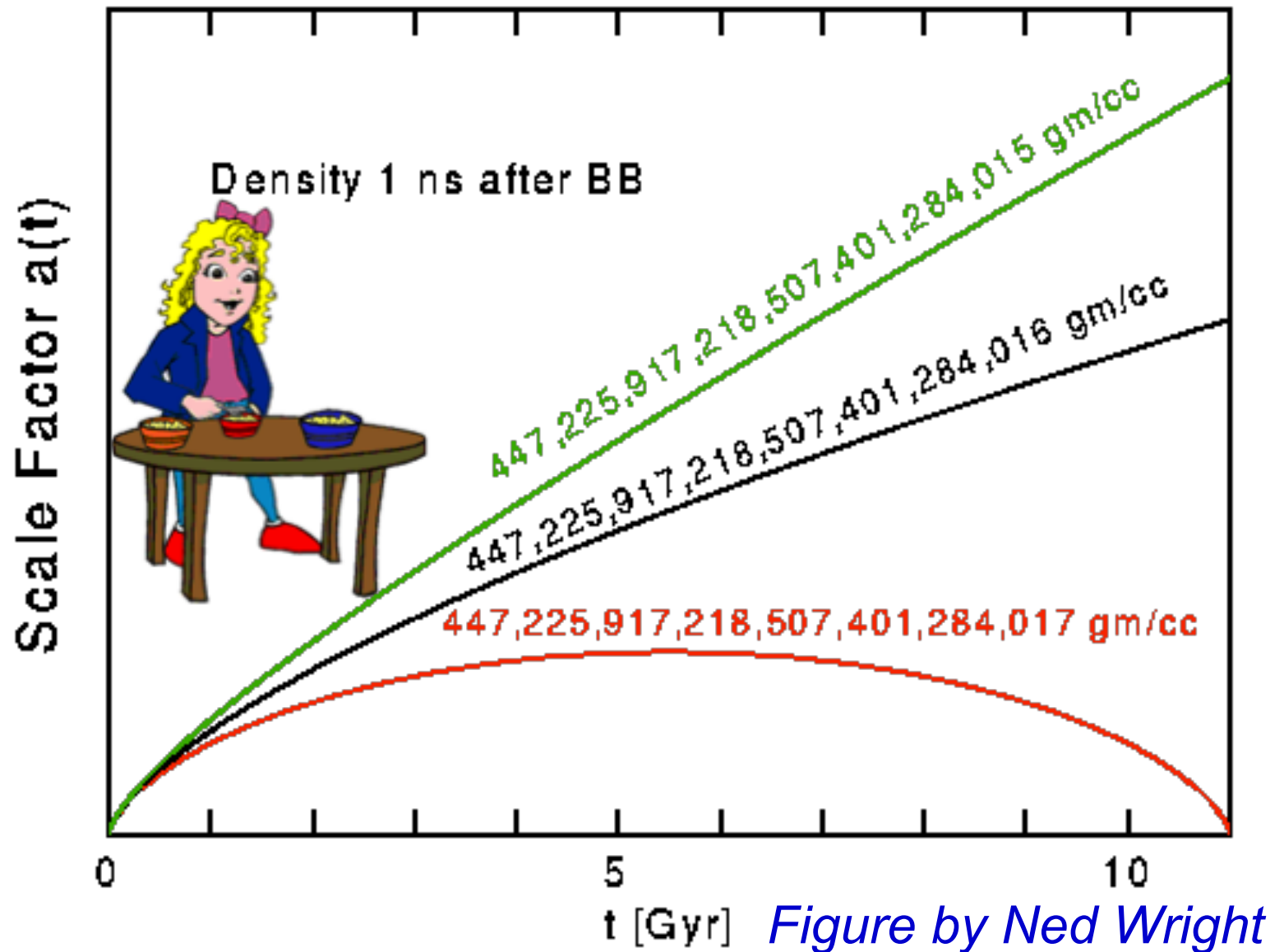


The "Horizon Problem":



How did causally disconnected regions know to bang at the same time?

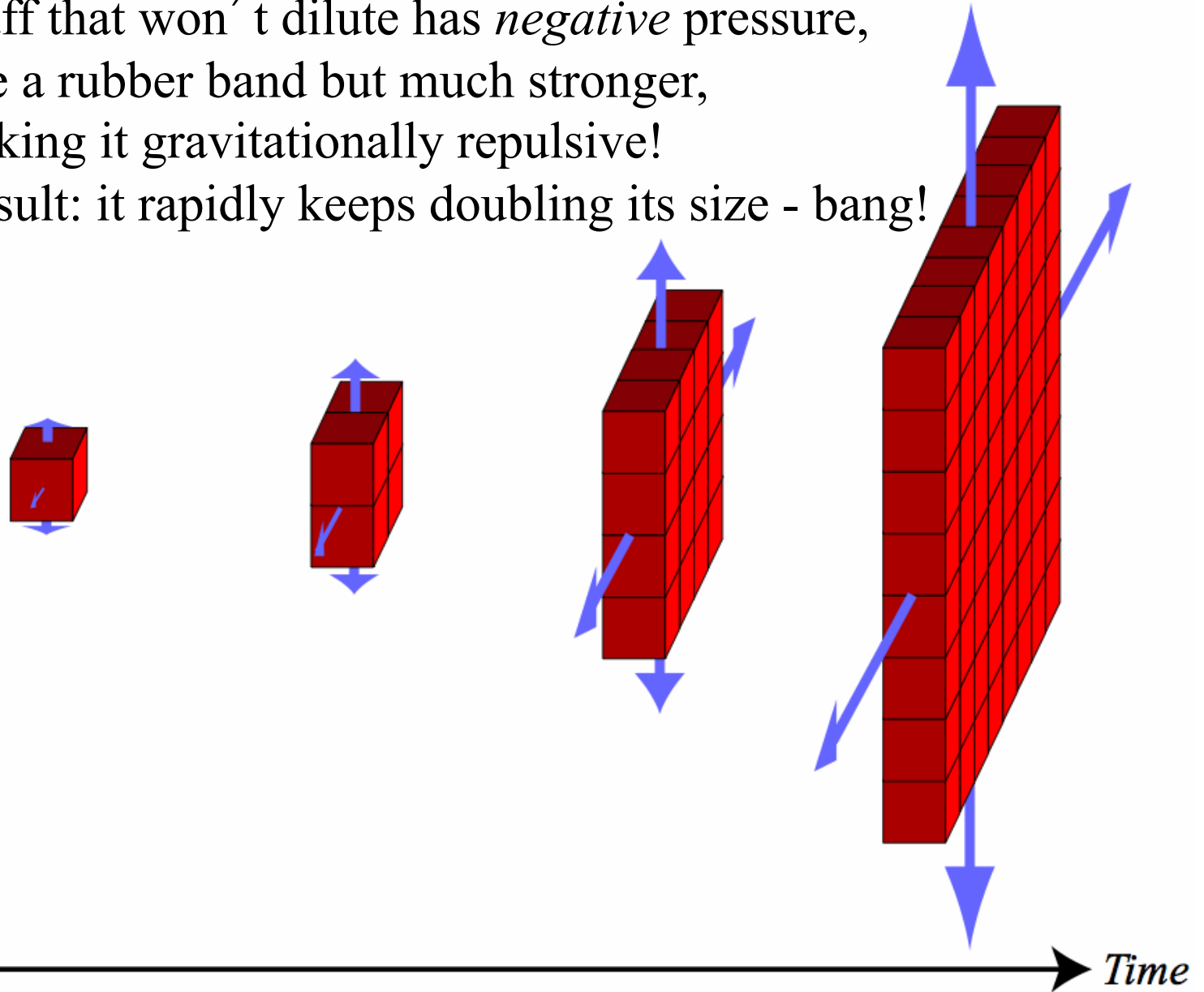
The “Flatness Problem”: Why was our Big Bang so “Goldilocks”?





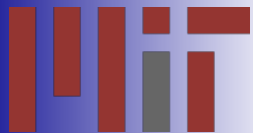
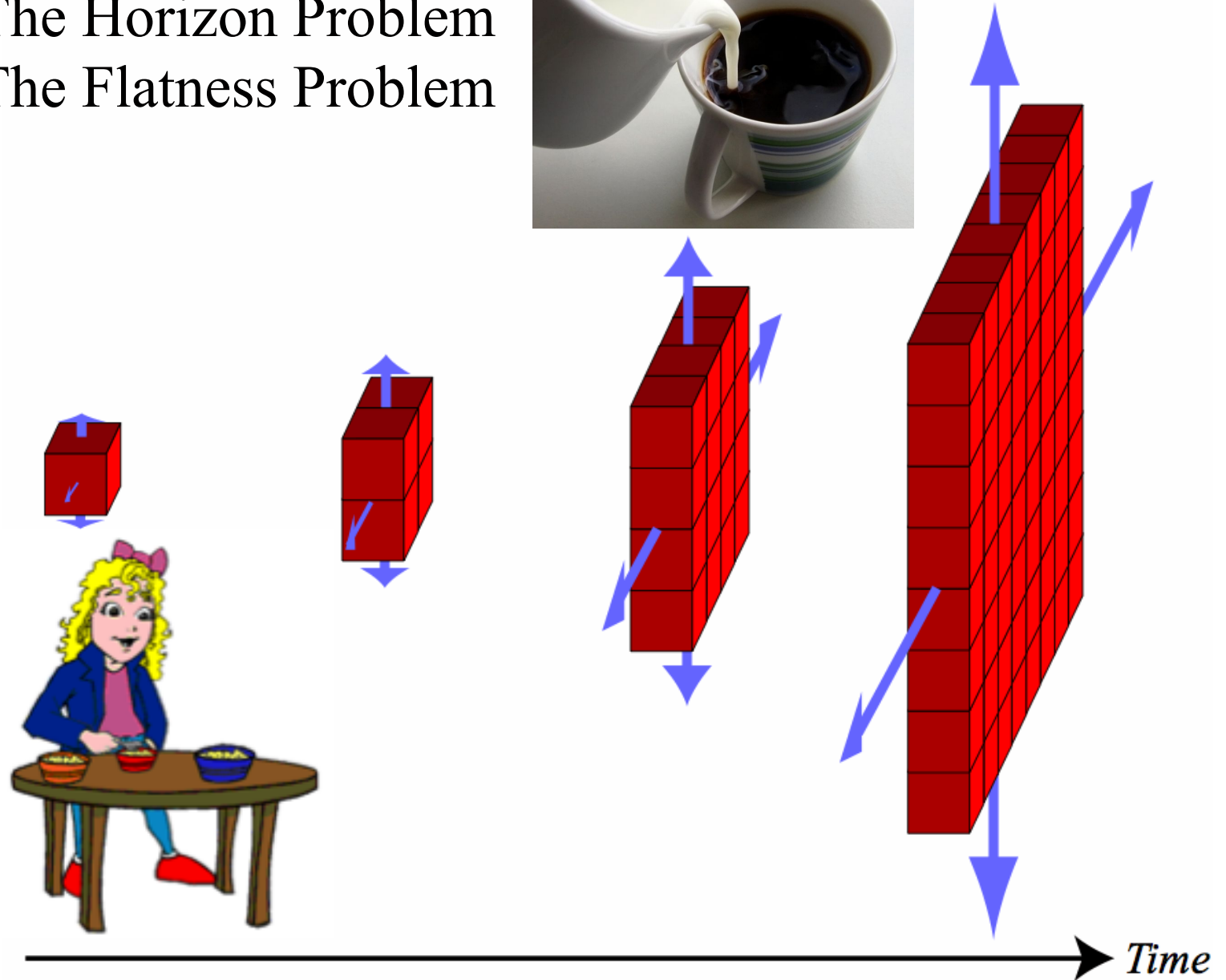
How inflation works:

- Einstein: source of gravity = density $+3p/c^2$
- Stuff that won't dilute has *negative* pressure, like a rubber band but much stronger, making it gravitationally repulsive!
- Result: it rapidly keeps doubling its size - bang!



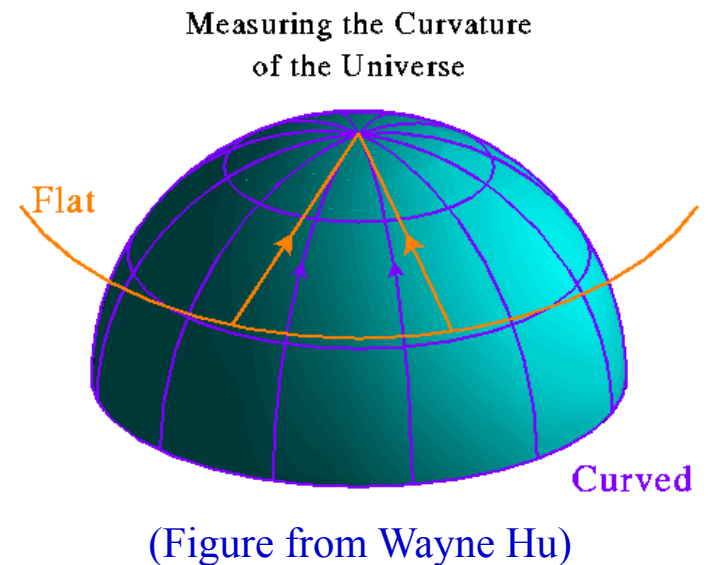
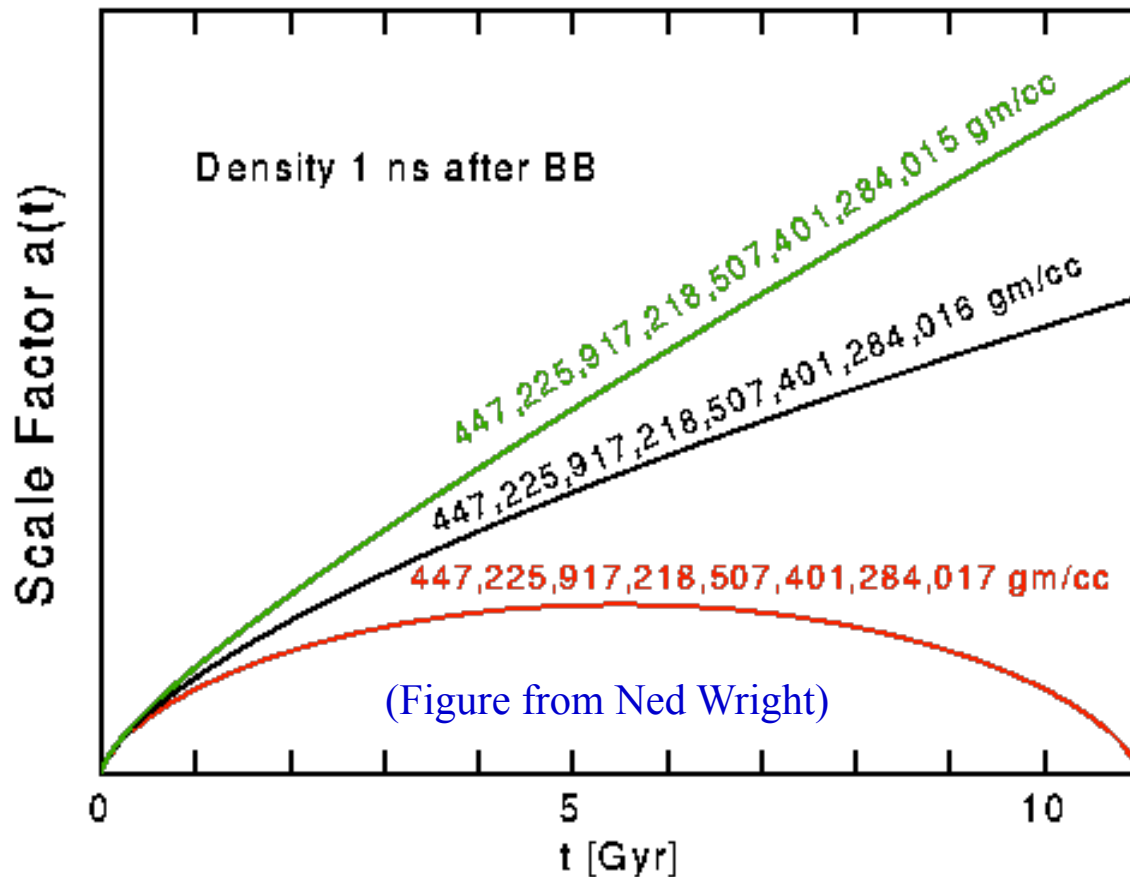
Inflation solves all three problems!

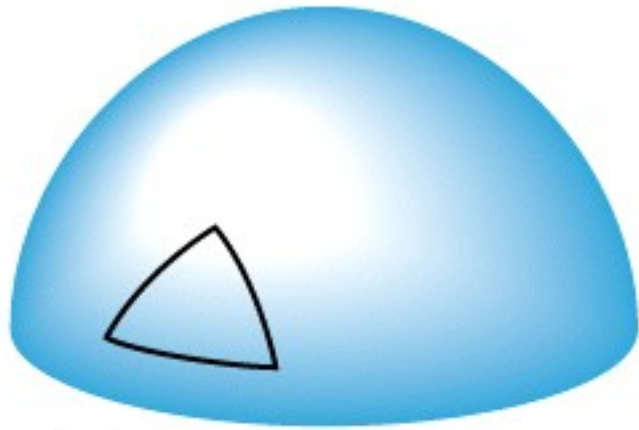
- The “Bang” Problem
- The Horizon Problem
- The Flatness Problem



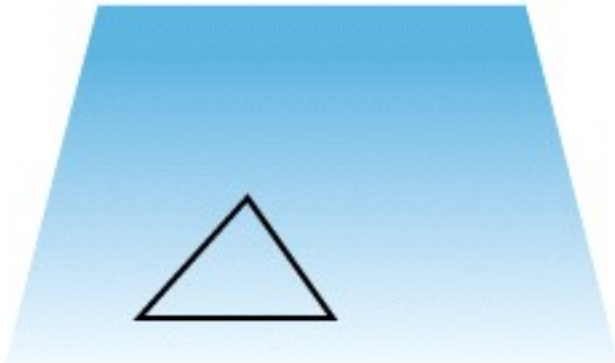
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Evidence #1 for inflation: Space is very flat

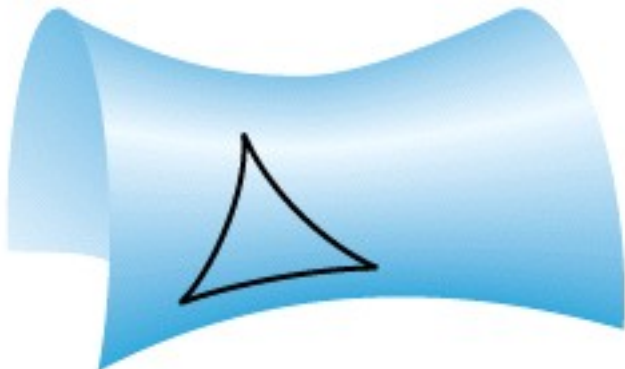




a Spherical space $\rho > \rho_c$



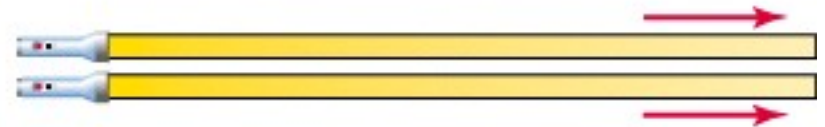
b Flat space $\rho = \rho_c$



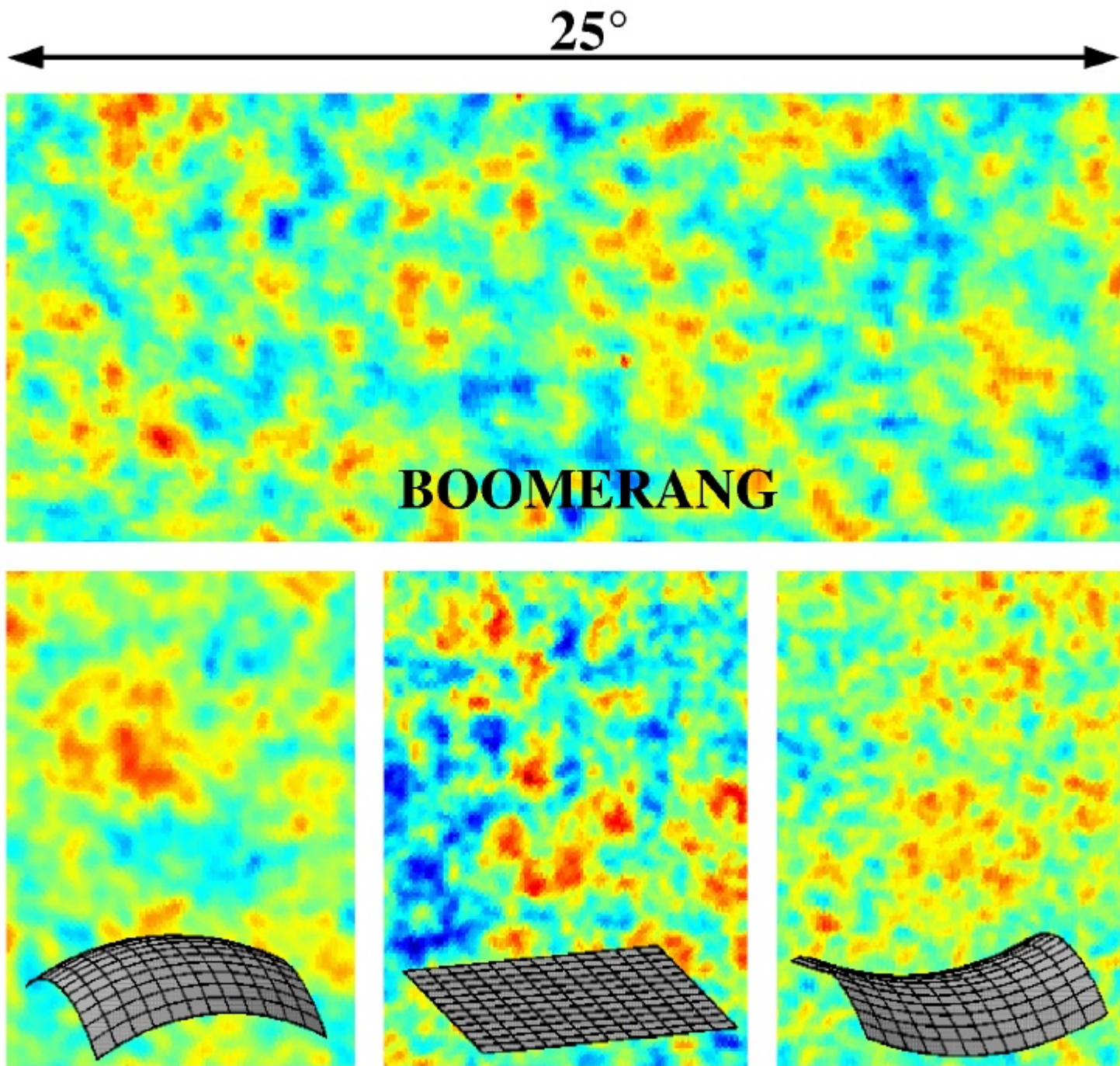
c Hyperbolic space $\rho < \rho_c$

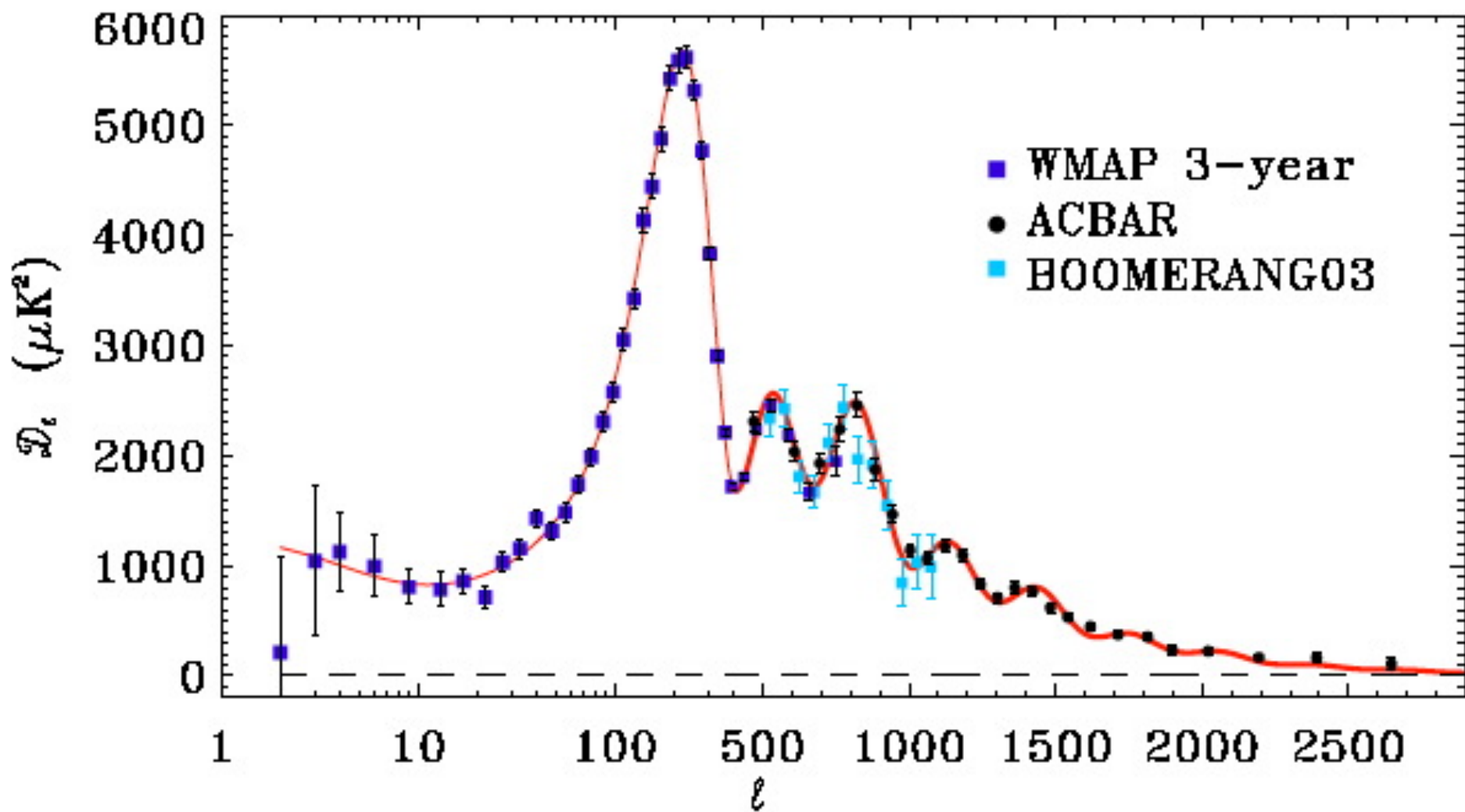


Alan
Guth



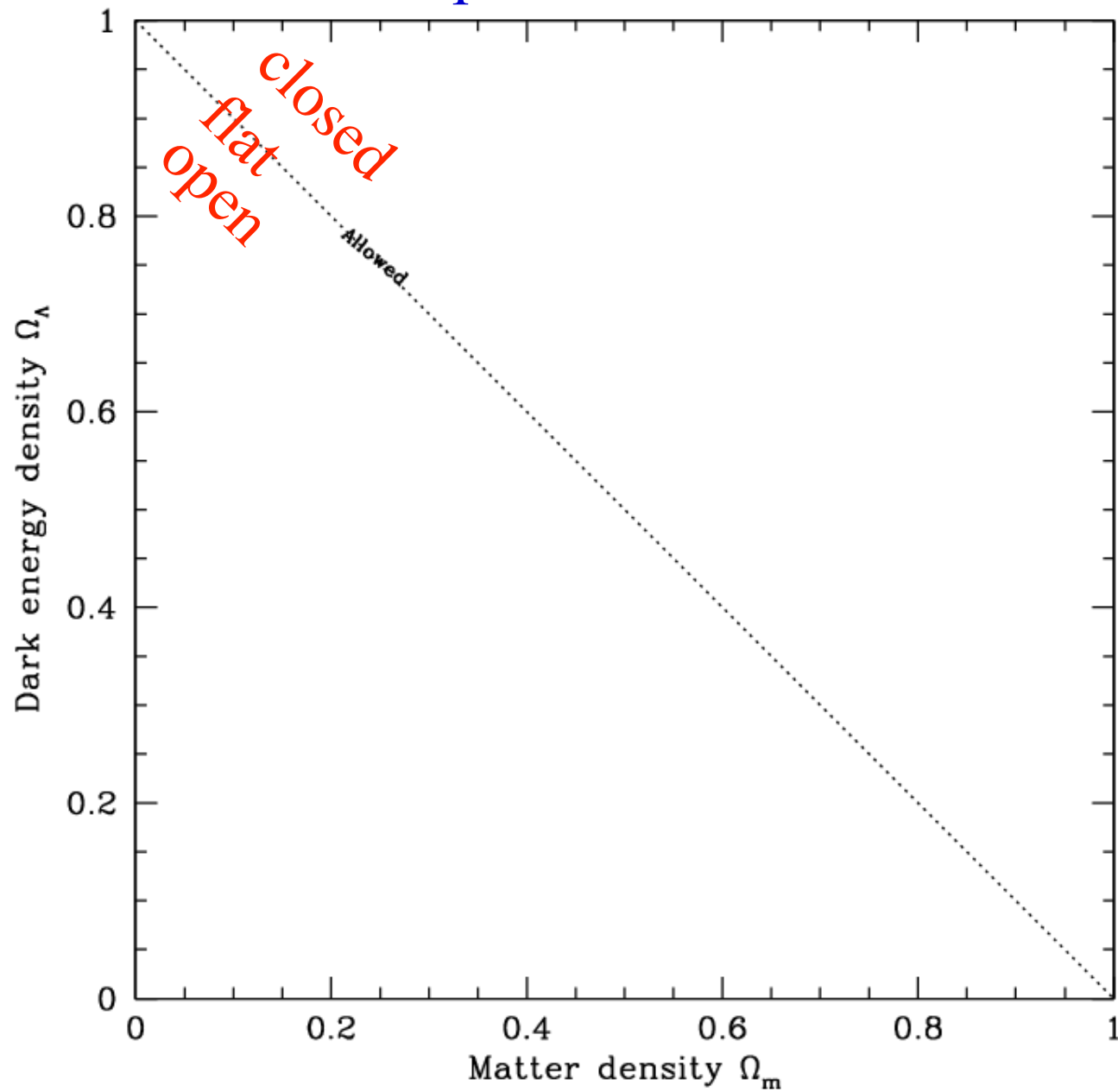
(Figure from Boomerang team)





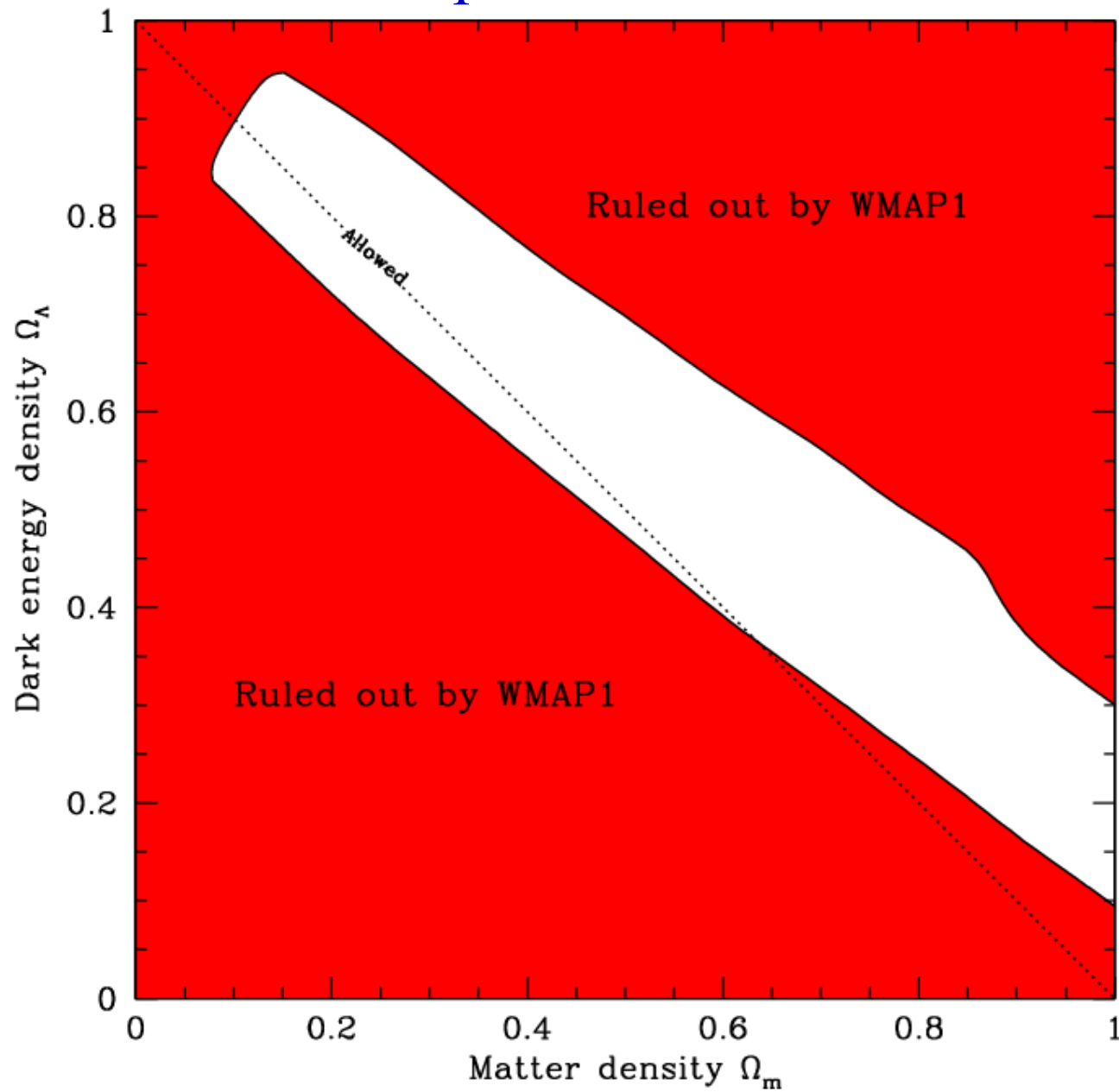
Reichardt et al 2008, arXiv:0801.1491

How flat is space?



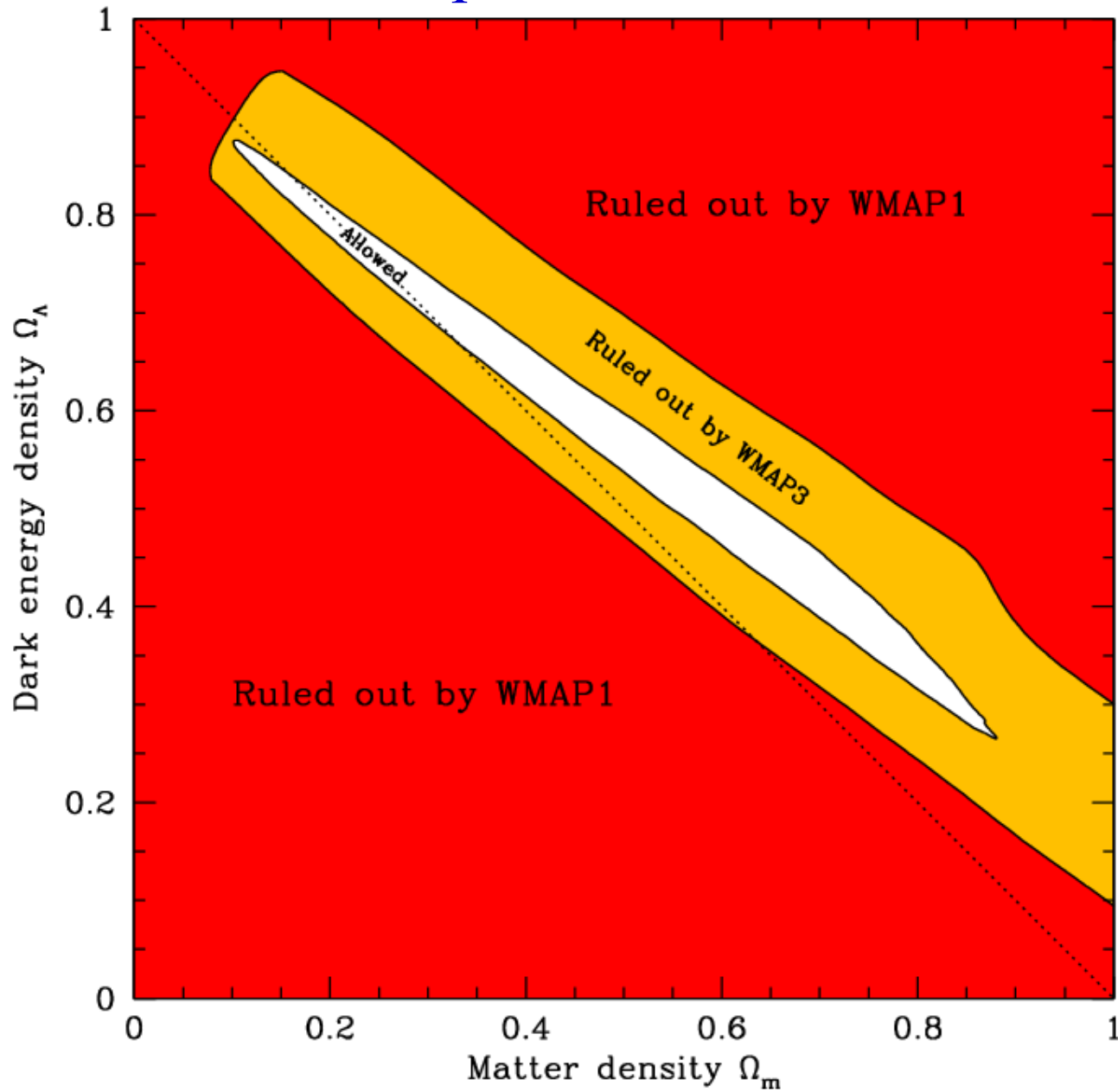
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How flat is space?



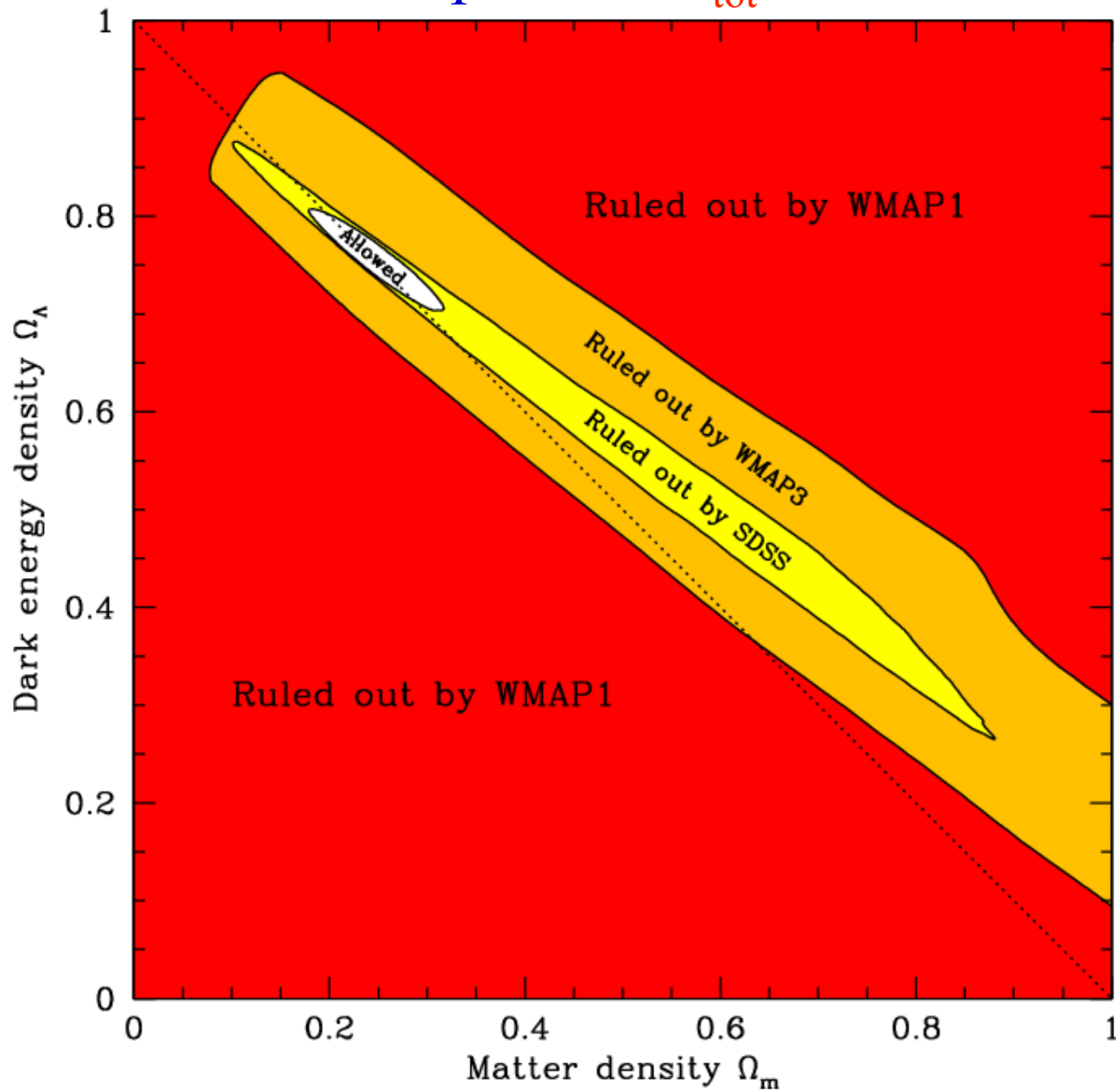
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How flat is space? Somewhat.



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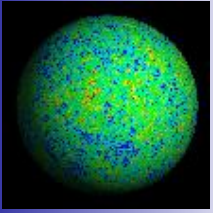
How flat is space? $\Omega_{\text{tot}} = 1.003 \pm 0.010$



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Party on!



CMB

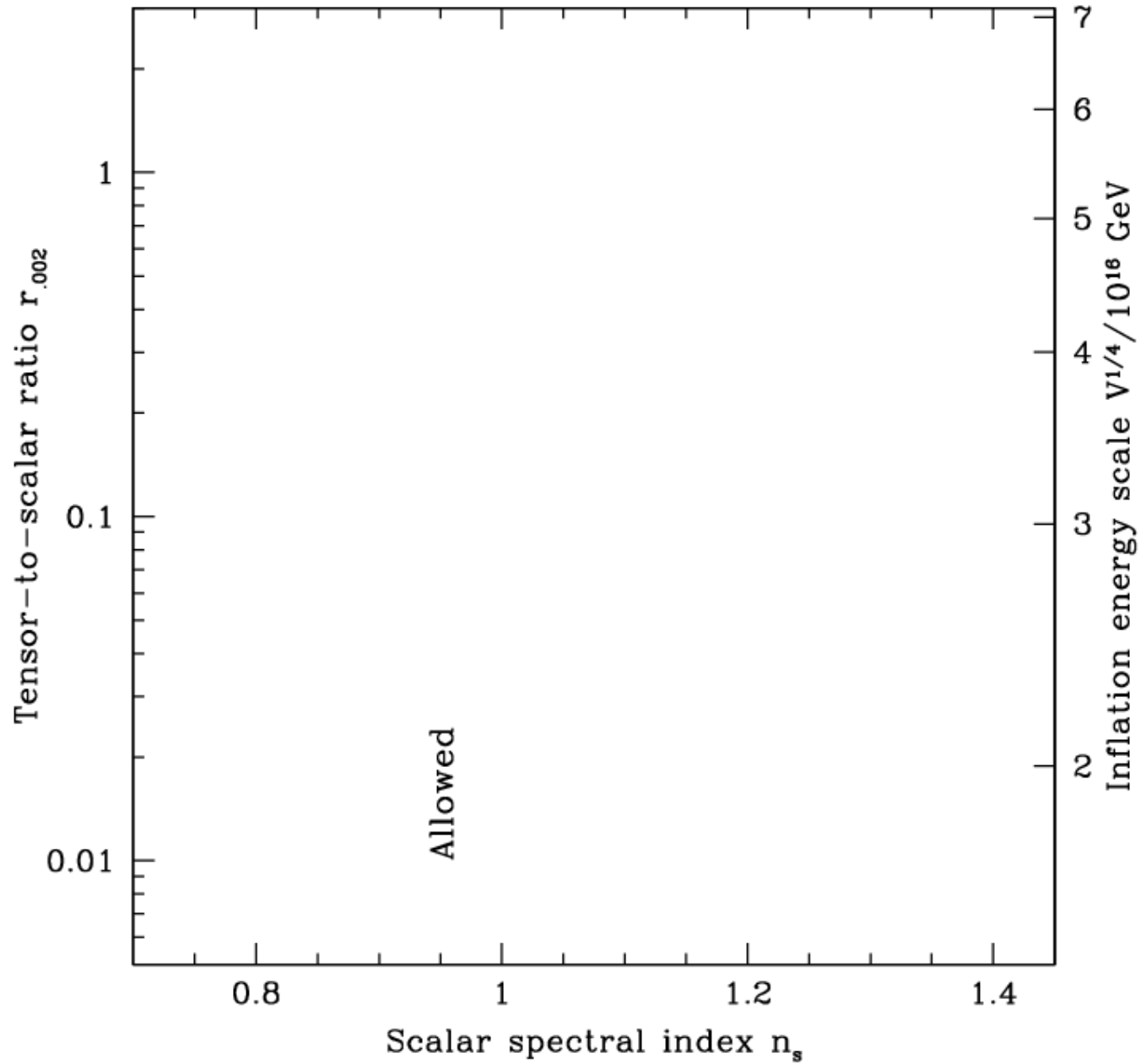
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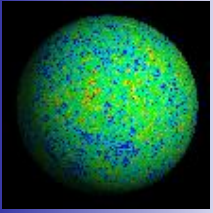


LSS



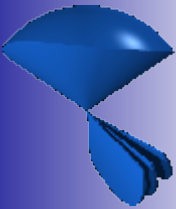
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CMB

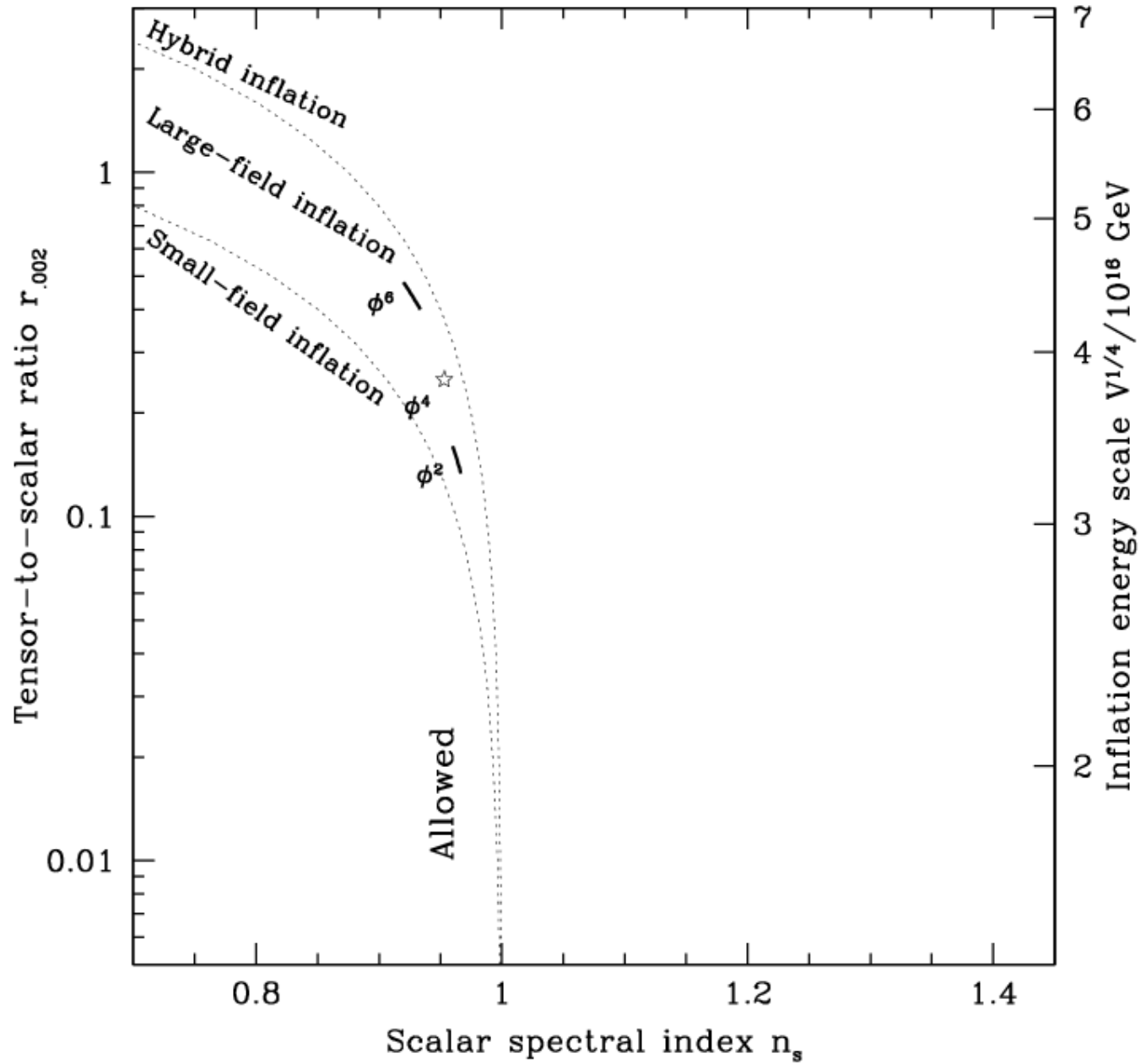
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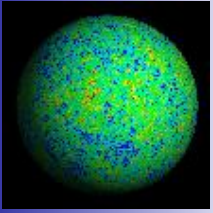


LSS



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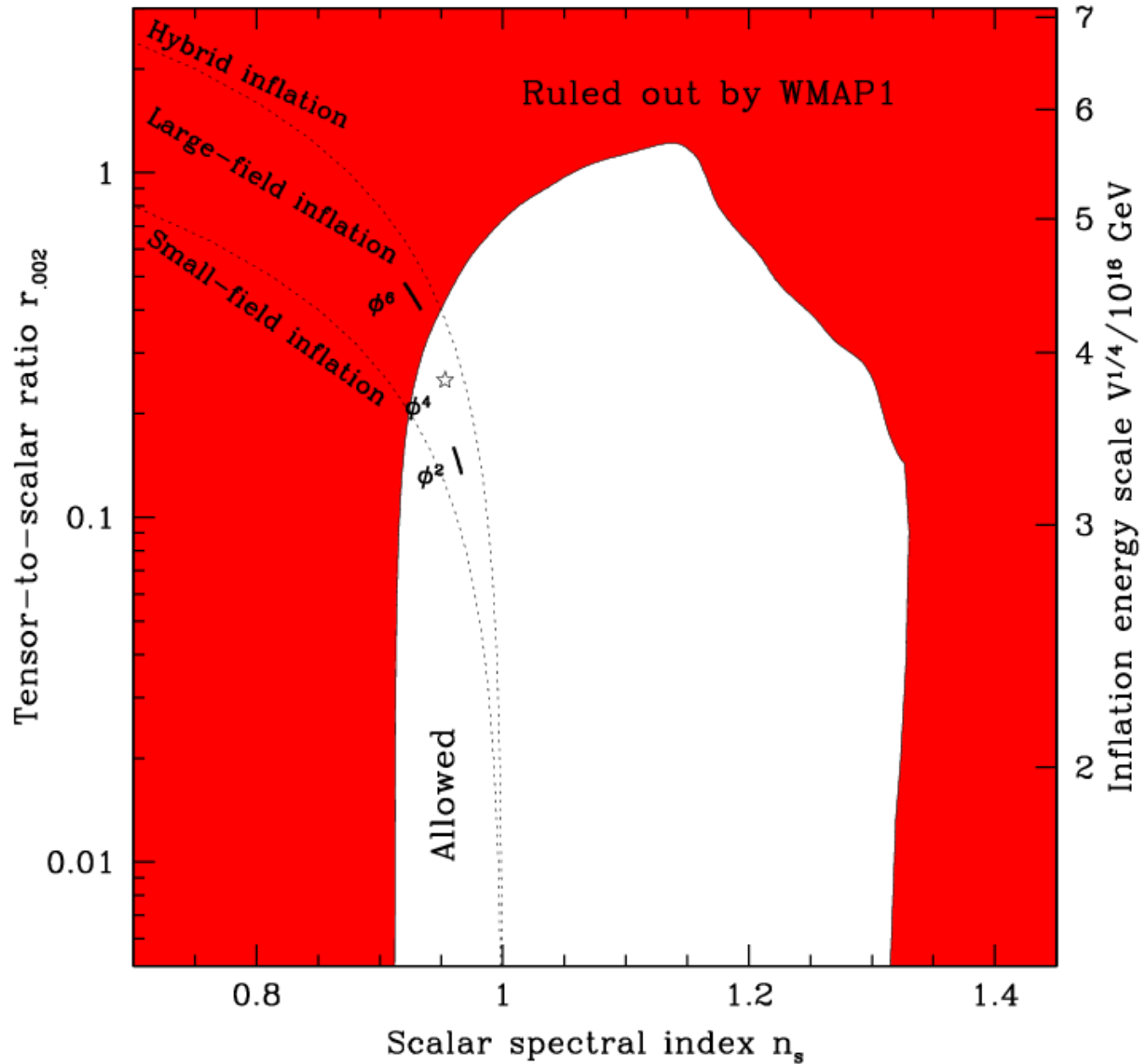
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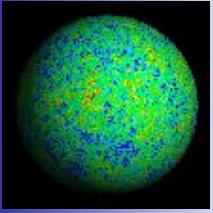


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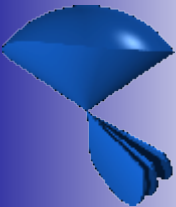
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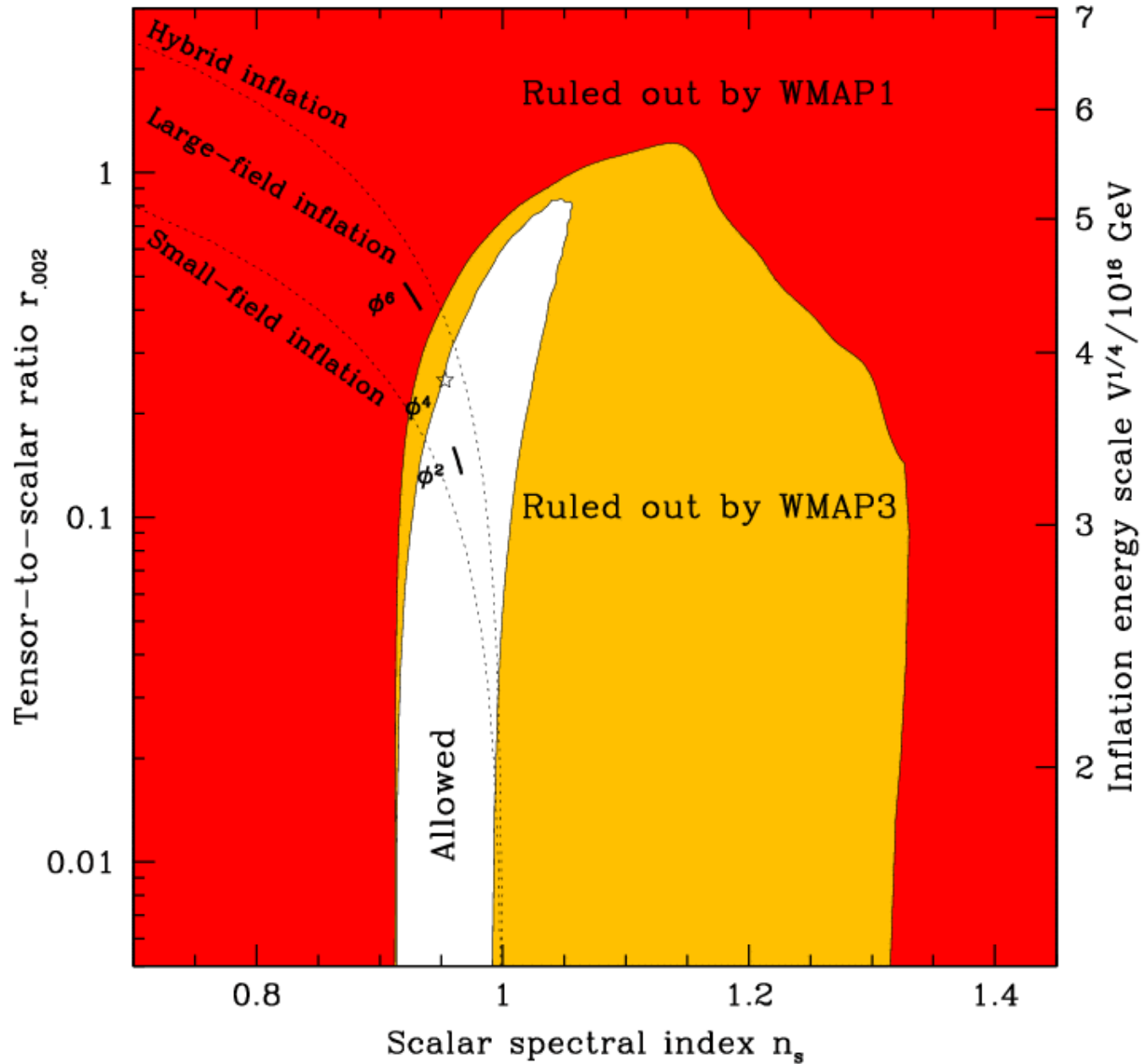
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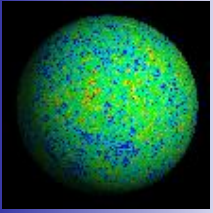


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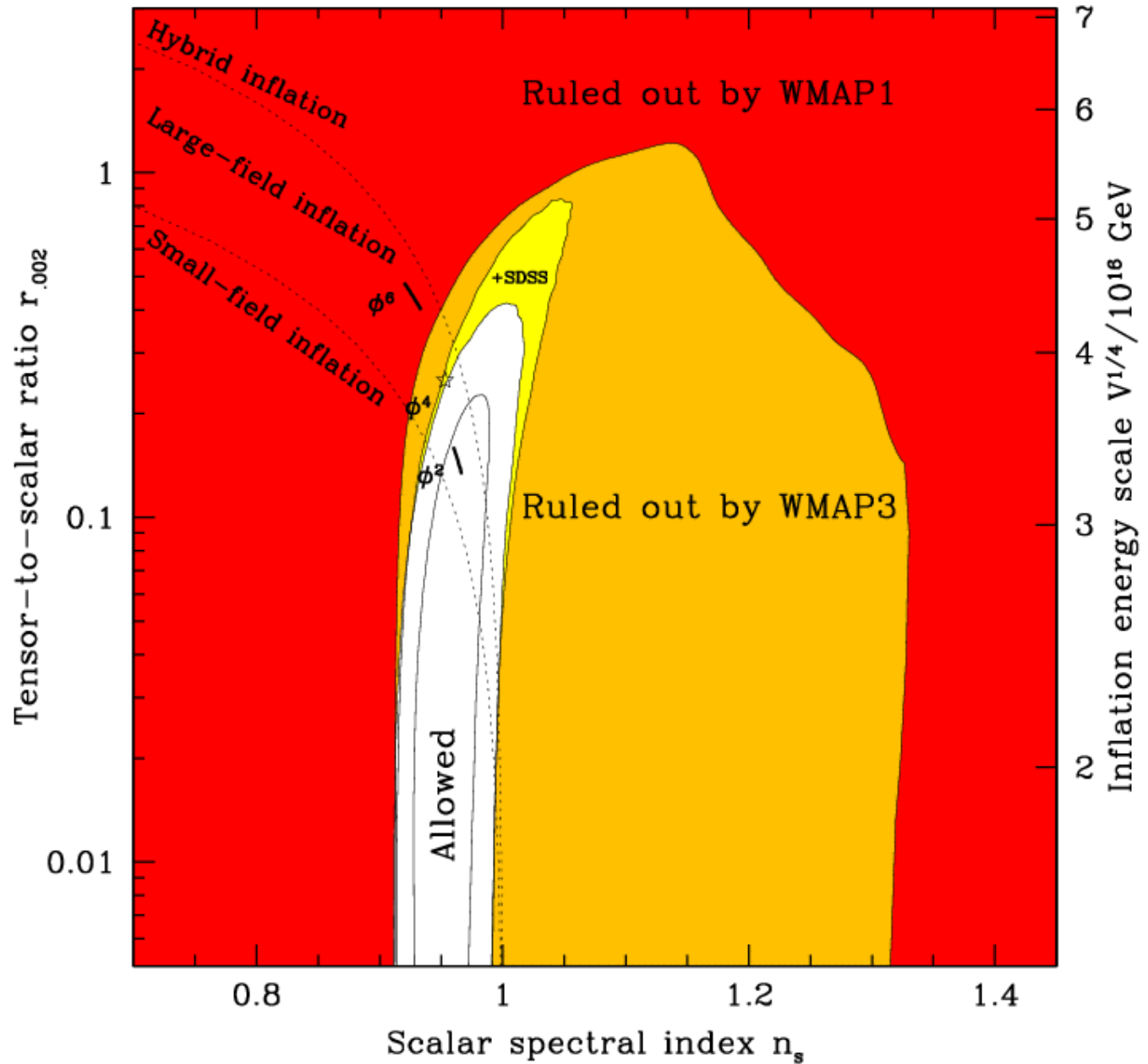
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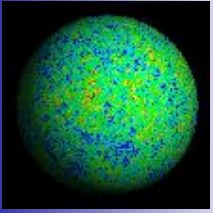


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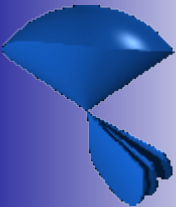
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CMB

+

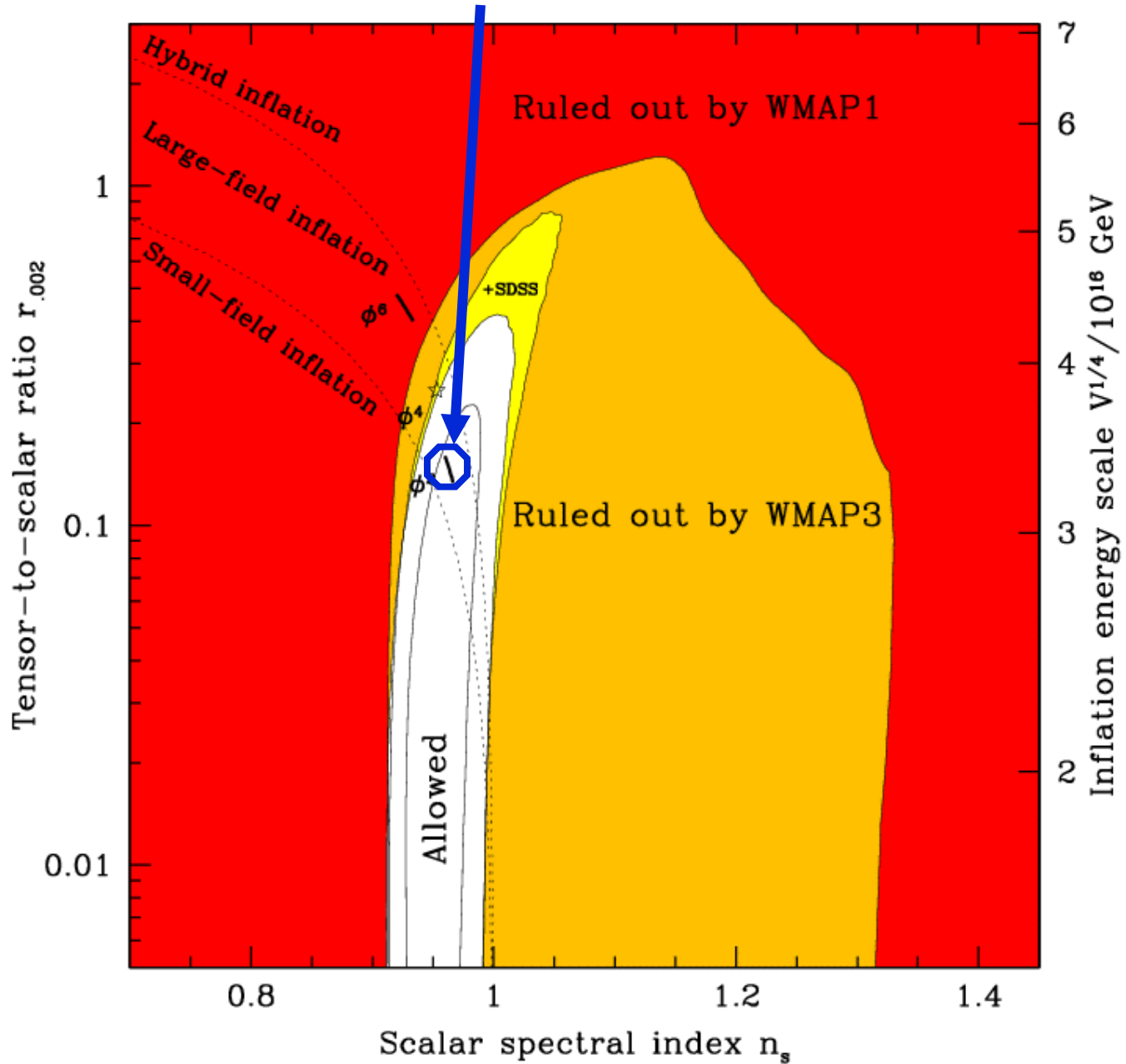


LSS



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Planck + SDSS: $\Delta n = 0.008$, $\Delta r = 0.012$



But be careful what
you wish for!

Inflation gave us
more than we'd
bargained for...

Parallel Universes



HEARING COLORS, TASTING SHAPES • ICEMAN REVISITED

SCIENTIFIC AMERICAN

MAY 2003
WWW.SCIAM.COM

Infinite Earths in

PARALLEL UNIVERSES

Really Exist

Orphan Drugs:
Too Successful?

Keys to Robust Networks

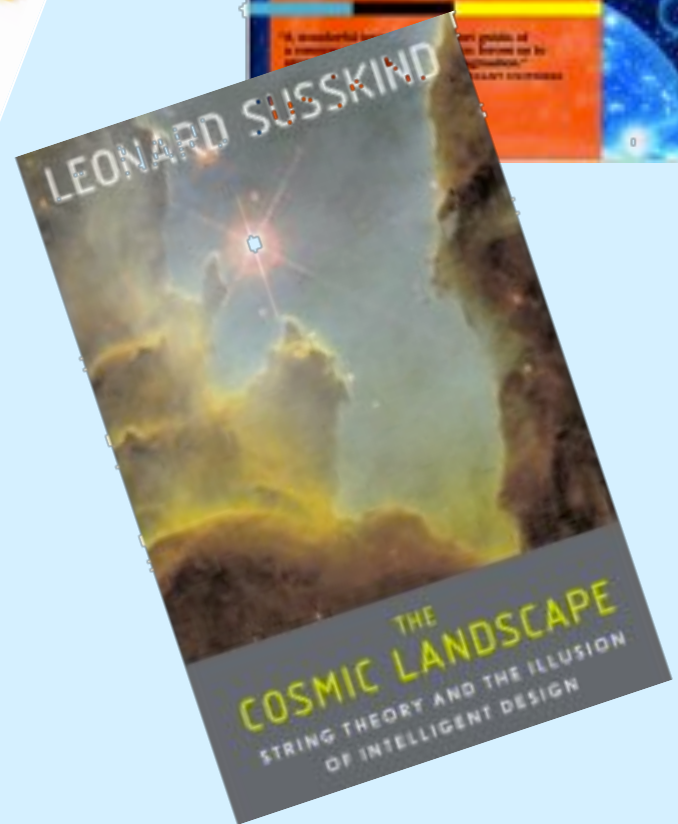
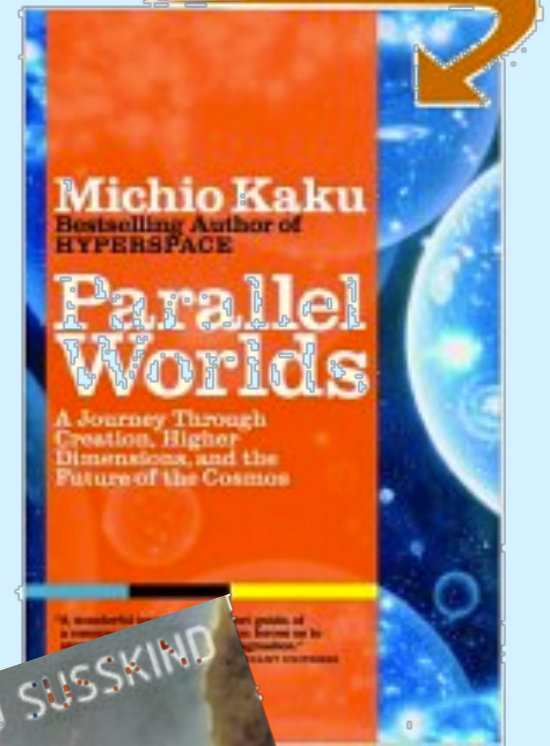
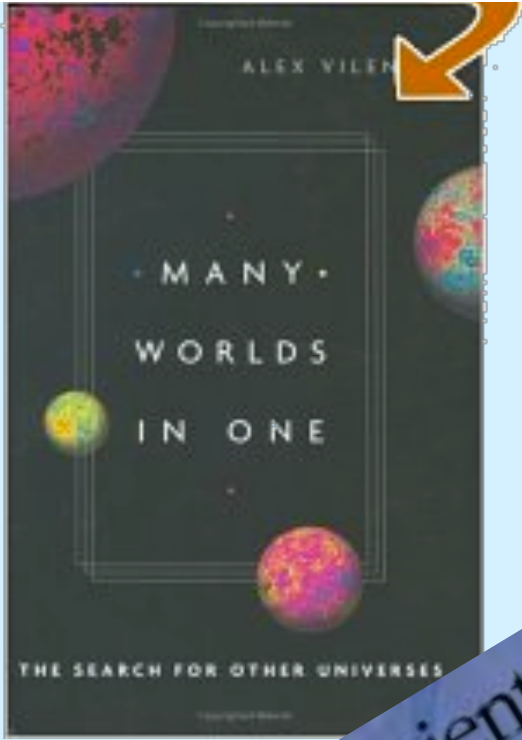
Smallpox Defense
Readiness

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\$4.95 U.S./\$6.50 CAN



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This isn't science!



It's inevitable



Makes sense!

Why not?

I hate it!



Party on!

Q: Is there more that exists than we can see?



Cosmology suggests yes!



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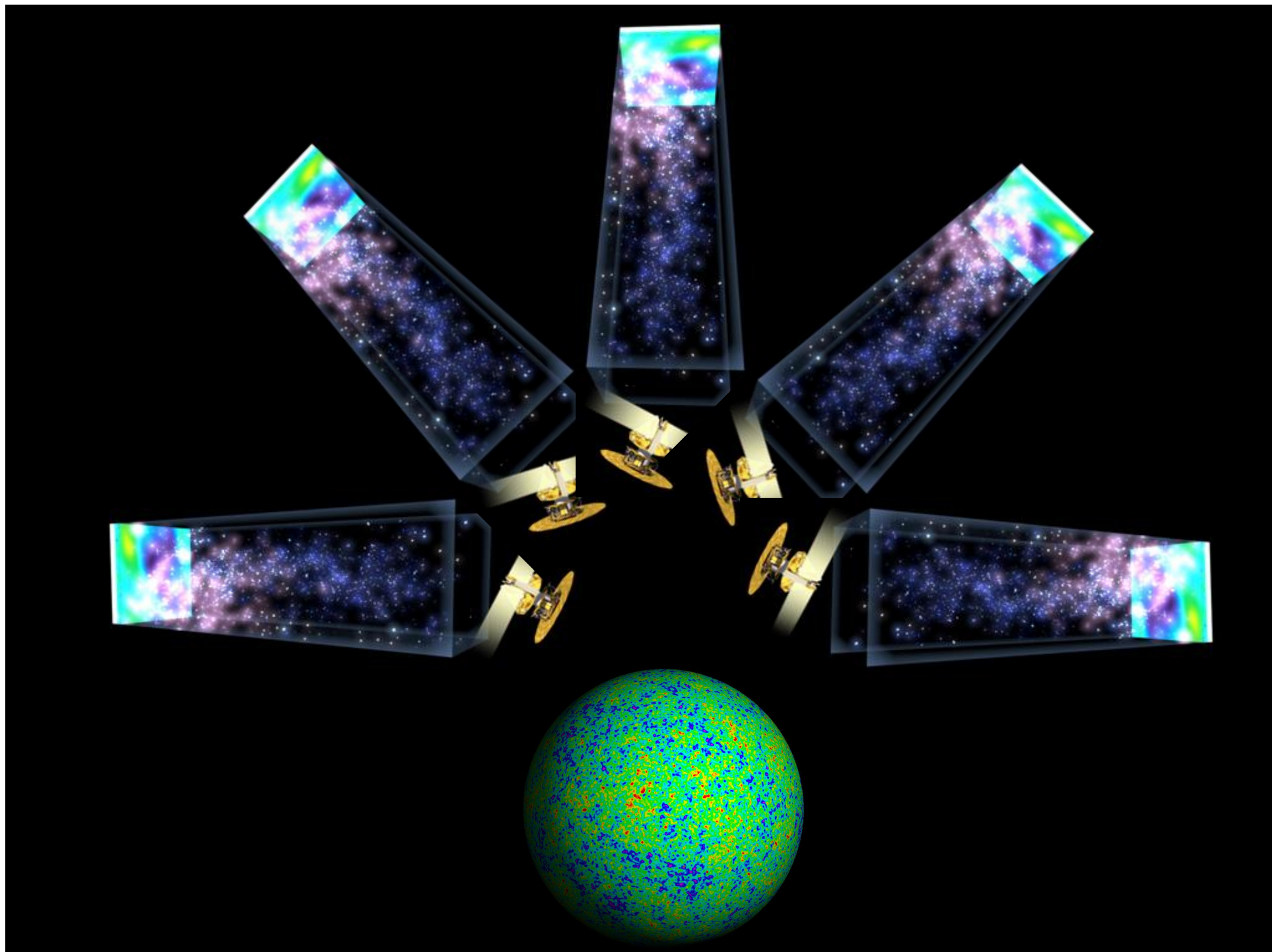
Outline:

- The most interesting question isn't whether parallel universes exist, but whether the multiverse has 1, 2, 3 or 4 levels
- Evidence for parallel universes
- Reasons to like/dislike multiverses

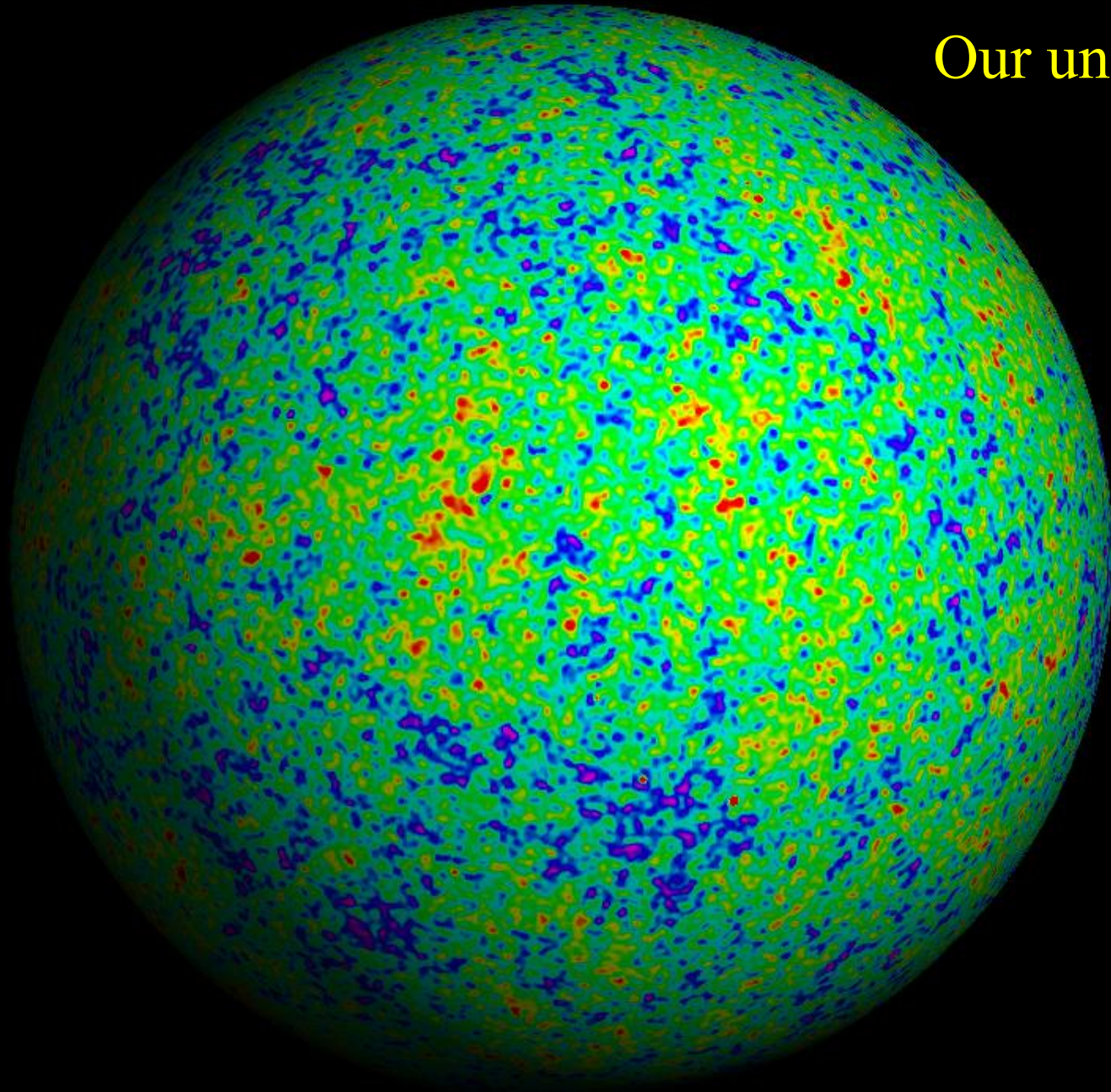


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What do we mean
by *our* universe?



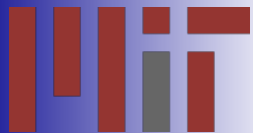
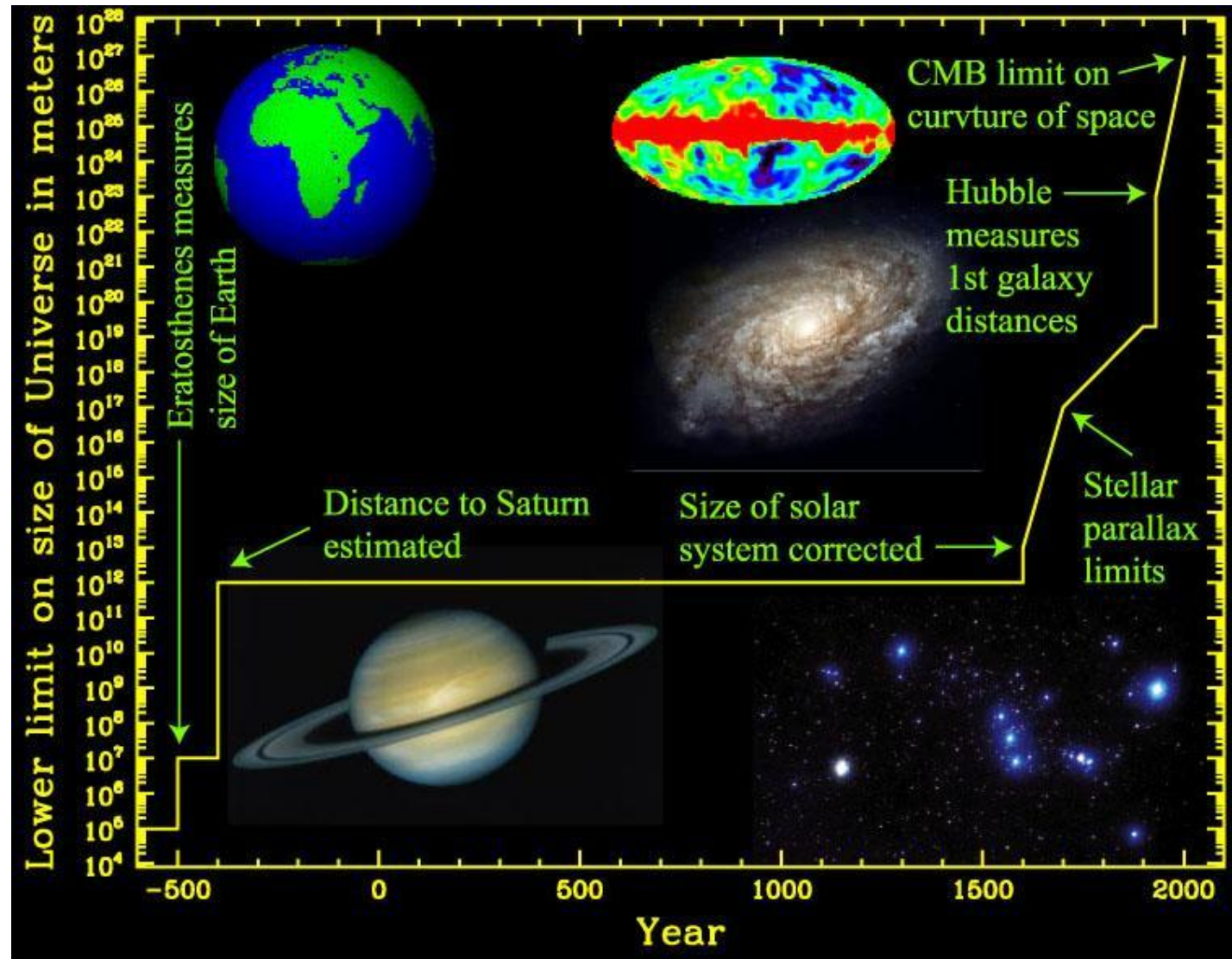
Foreground-cleaned WMAP map from Tegmark, de Oliveira-Costa & Hamilton, astro-ph/0302496



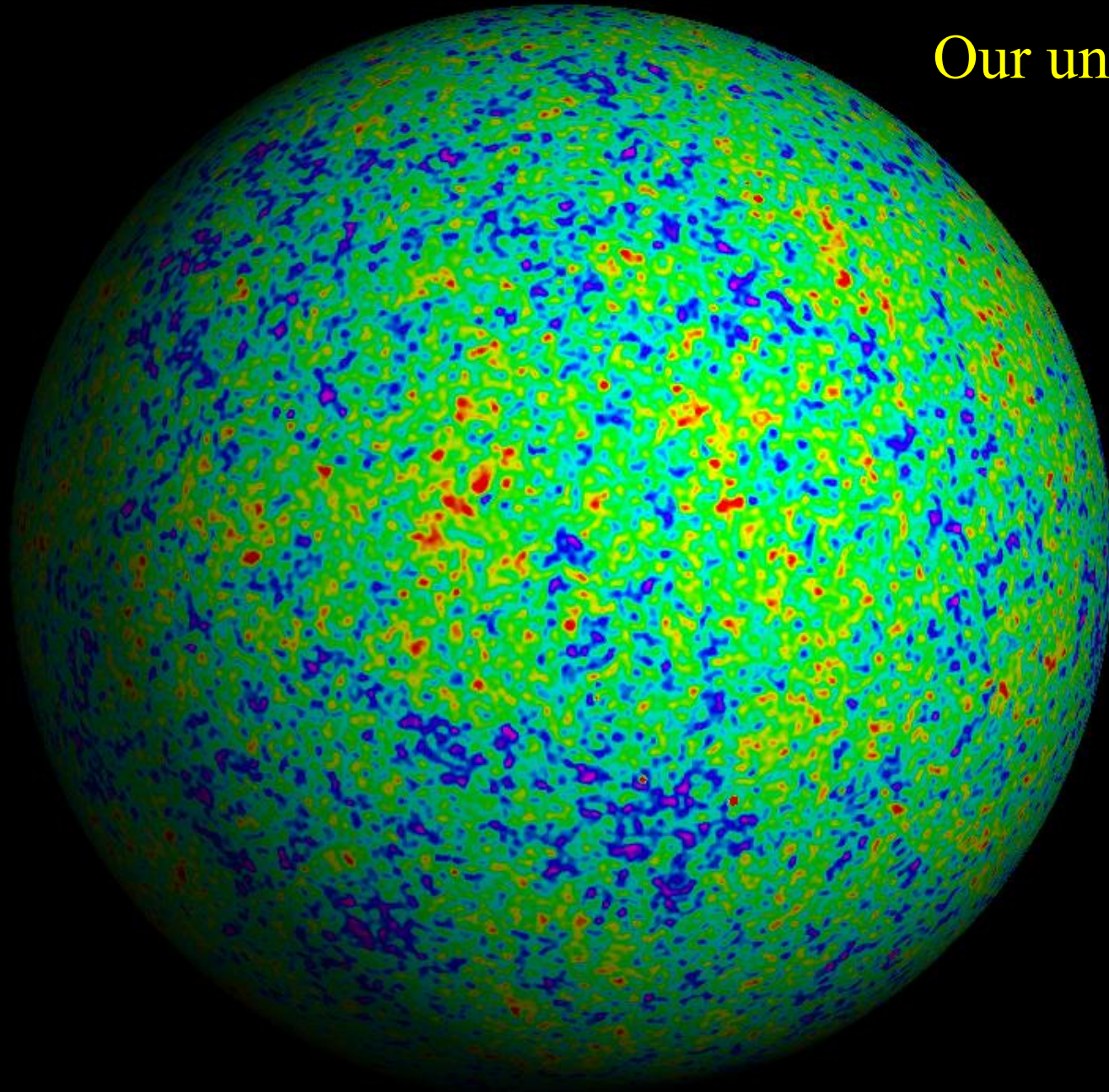
Our universe

Level I

How big is our space?



Foreground-cleaned WMAP map from Tegmark, de Oliveira-Costa & Hamilton, astro-ph/0302496



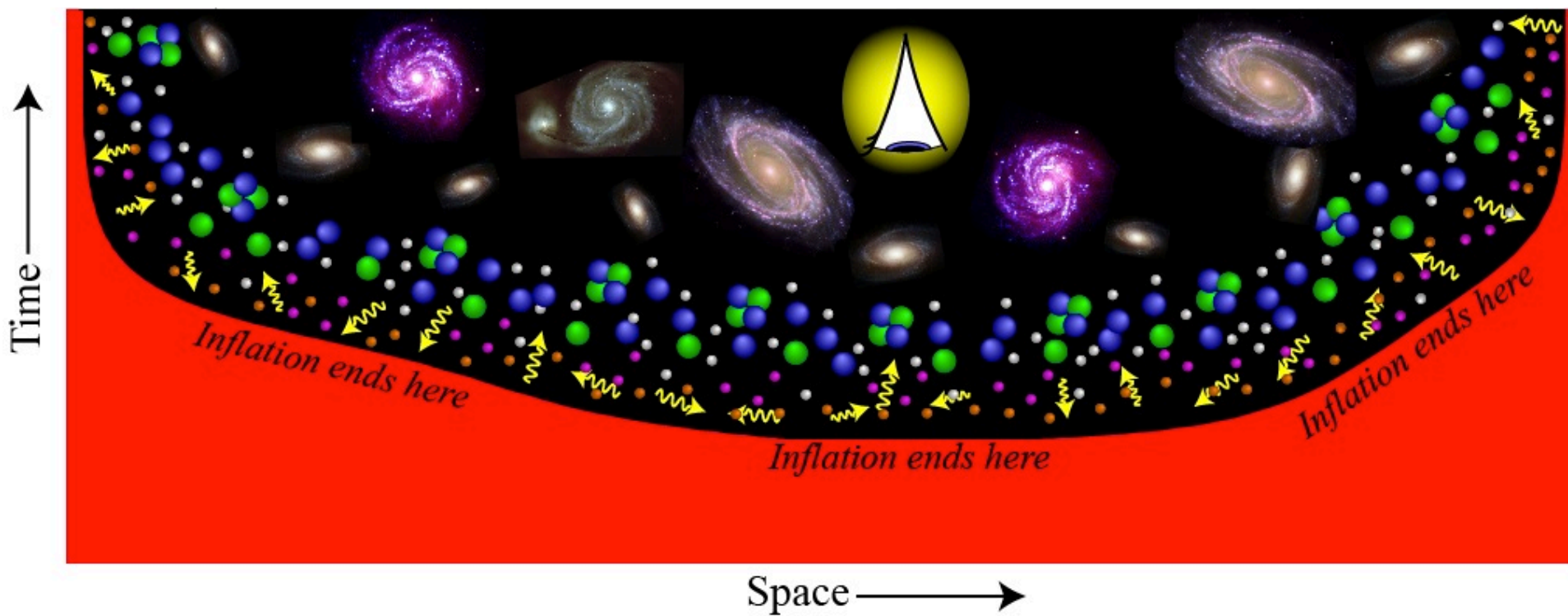
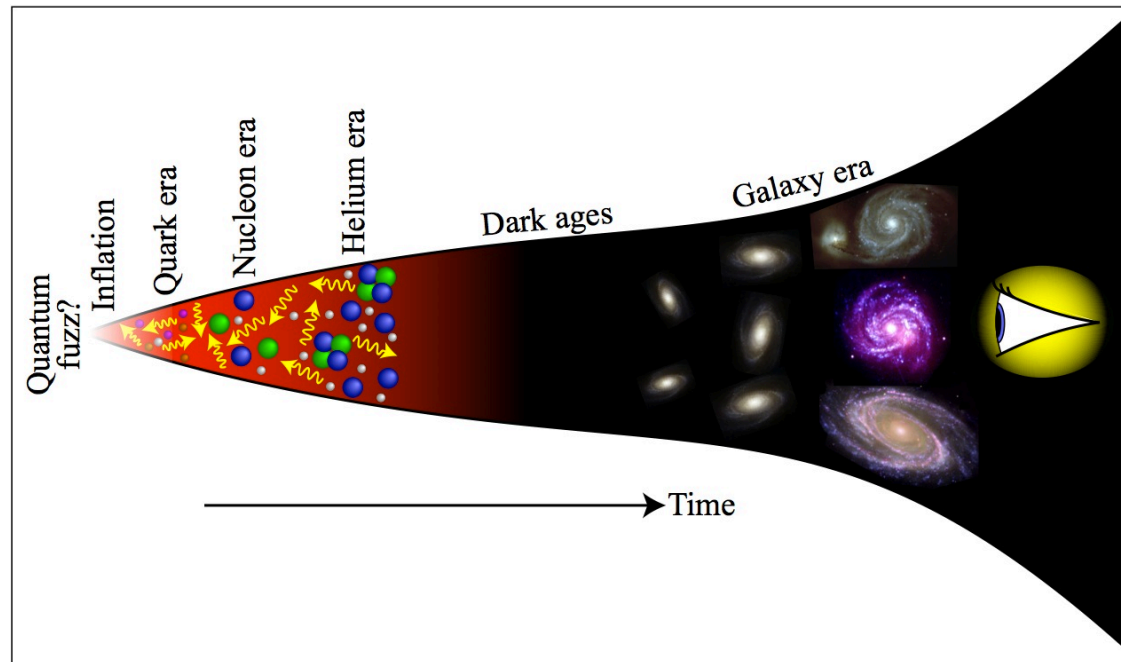
Our universe

Inflation generically predicts that space isn't just huge, but truly infinite!



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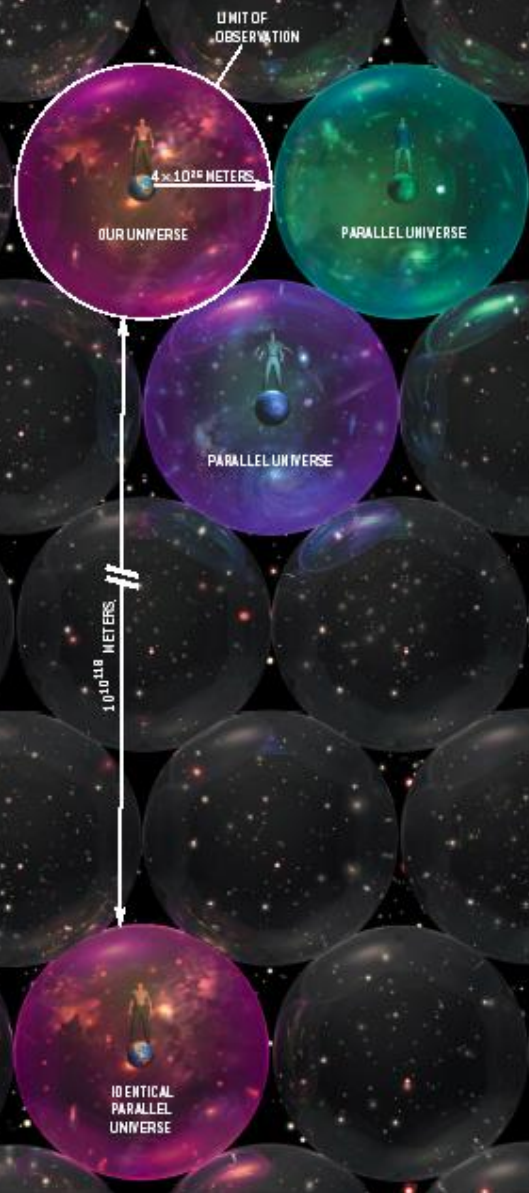
How to make an infinite space inside a finite volume



LEVEL I MULTIVERSE

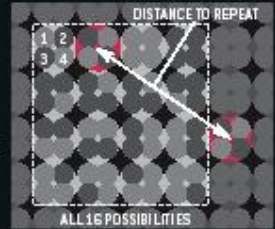
THE SIMPLEST TYPE of parallel universe is simply a region of space that is too far away for us to have seen yet. The farthest that we can observe is currently about 4×10^{26} meters, or 42 billion light-years—the distance that light has been able to travel since the big

bang began. (The distance is greater than 14 billion light-years because cosmic expansion has lengthened distances.) Each of the Level I parallel universes is basically the same as ours. All the differences stem from variations in the initial arrangement of matter.



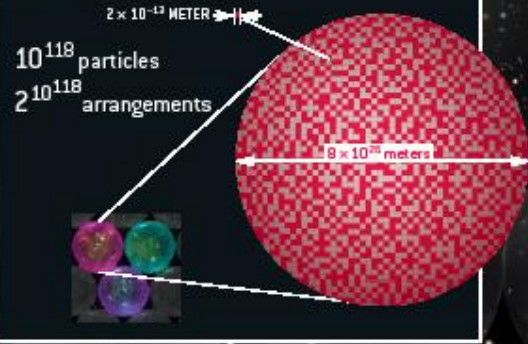
How Far Away is a Duplicate Universe?

EXAMPLE UNIVERSE
 Imagine a two-dimensional universe with space for four particles. Such a universe has 2^4 , or 16, possible arrangements of matter. If more than 16 of these universes exist, they must begin to repeat. In this example, the distance to the nearest duplicate is roughly four times the diameter of each universe.



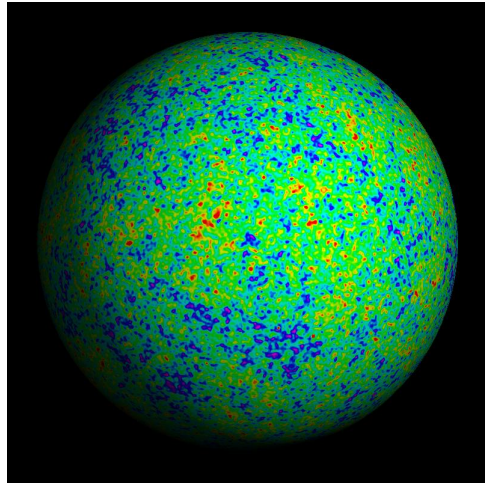
4 particles
 2^4 arrangements

OUR UNIVERSE
 The same argument applies to our universe, which has space for about 10^{118} subatomic particles. The number of possible arrangements is therefore 2 to the 10^{118} , or approximately 10 to the 10^{118} . Multiplying by the diameter of the universe gives an average distance to the nearest duplicate of 10 to the 10^{118} meters.



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Multiverse level I: other Hubble volumes beyond our cosmic horizon



Giordano Bruno (executed 1600)
Ellis & Brundrit 1979, Q.J.R. Astr. Soc. 20, 37
Garriga & Vilenkin 2001, Phys.Rev. D64, 043511

Features:

- Same (effective) laws of physics
- Different initial conditions

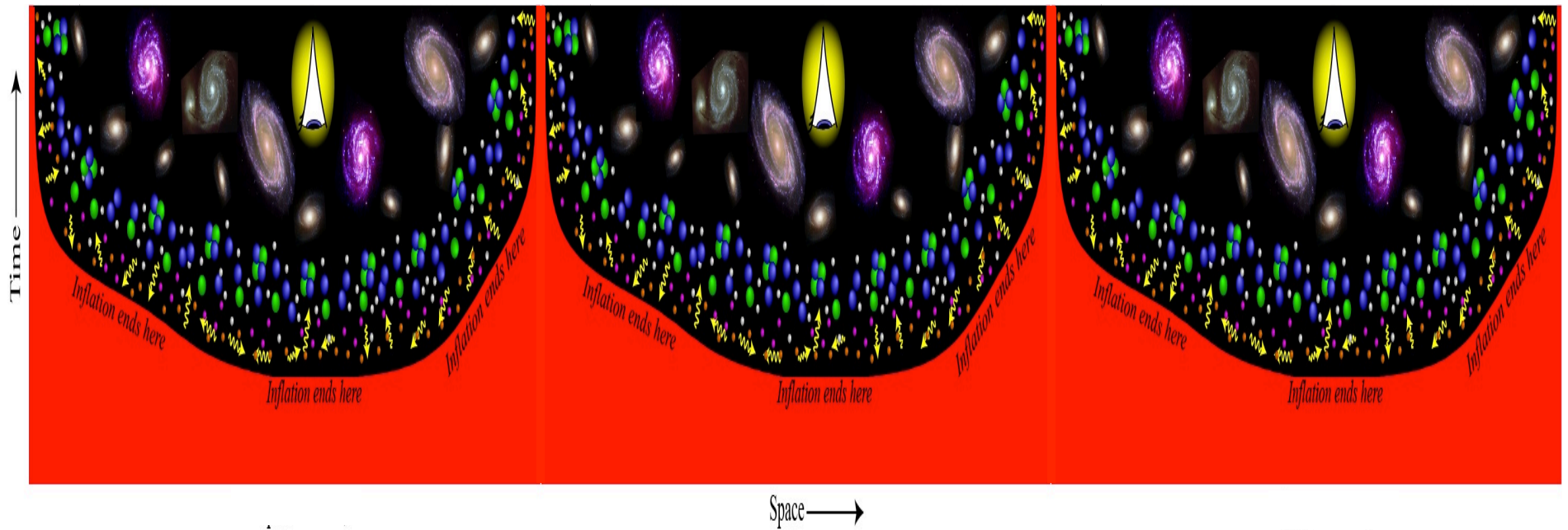
For our flat “concordance” cosmological model:

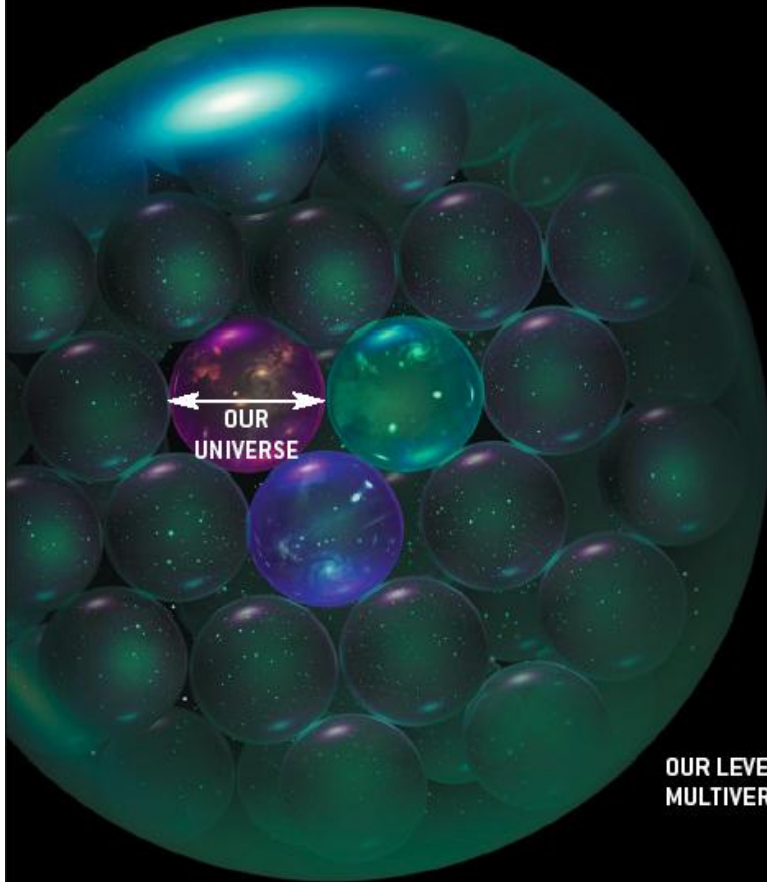
- Size of our Hubble volume $\sim 10^{26}$ m,
- Closest copy of you $\sim 10^{10^{29}}$ m
- Closest 100 lightyear bubble like ours $\sim 10^{10^{91}}$ m
- Closest Hubble volume like ours $\sim 10^{10^{118}}$ m



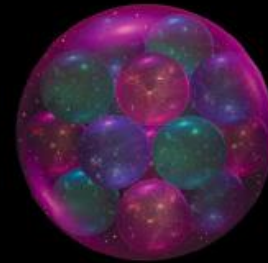
Level III

Level II Multiverse



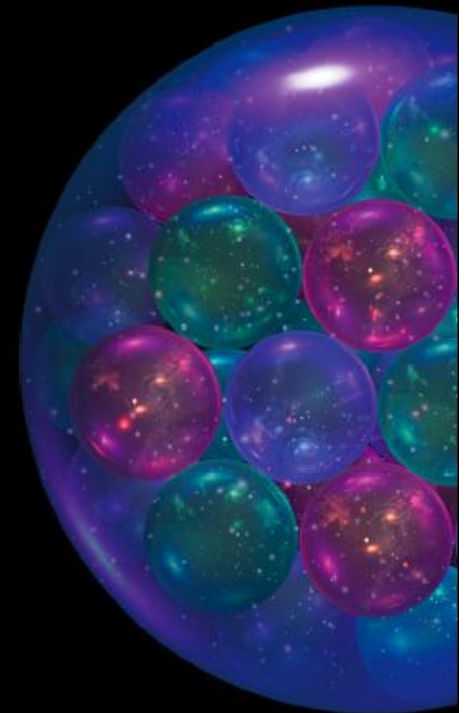


OUR LEVEL I
MULTIVERSE



PARALLEL
LEVEL I
MULTIVERSE

EMPTY
SPACE
(INFLATING)



PARALLEL
LEVEL I
MULTIVERSE

Why these values?

Standard model parameters:

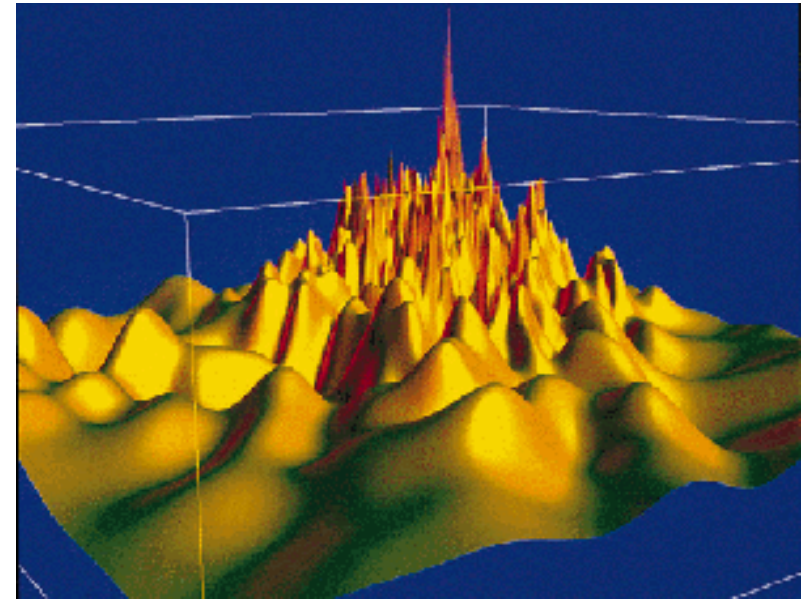
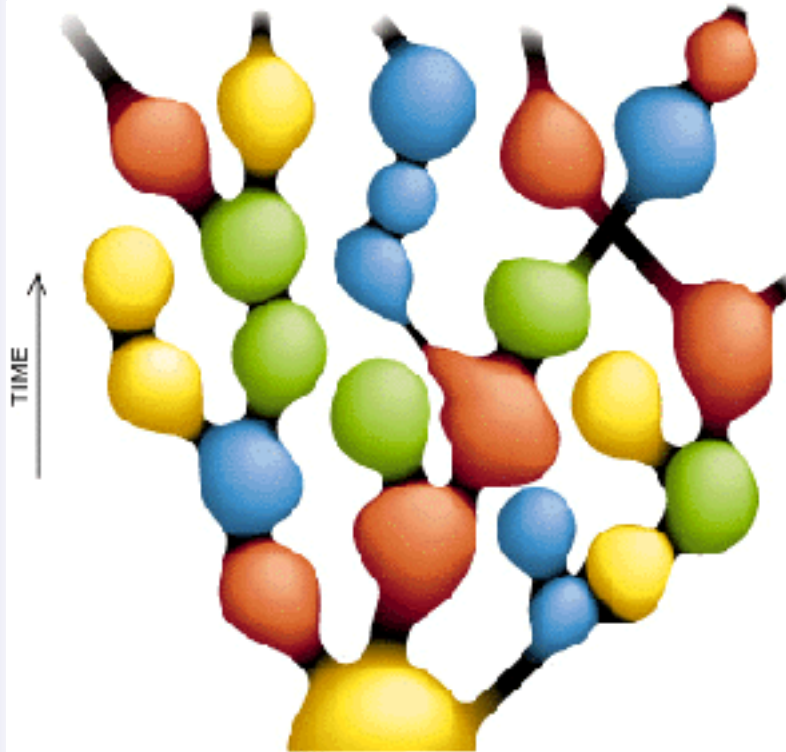
Particle physics

Cosmology

Parameter	Meaning	Measured value
g	Weak coupling constant at m_Z	0.6520 ± 0.0001
θ_W	Weinberg angle	0.48290 ± 0.00005
g_s	Strong coupling constant at m_Z	1.221 ± 0.022
μ^2	Quadratic Higgs coefficient	$\sim -10^{-33}$
λ	Quartic Higgs coefficient	$\sim 1?$
G_e	Electron Yukawa coupling	2.94×10^{-6}
G_μ	Muon Yukawa coupling	0.000607
G_τ	Tauon Yukawa coupling	0.0102156233
G_u	Up quark Yukawa coupling	0.000016 ± 0.000007
G_d	Down quark Yukawa coupling	0.00003 ± 0.00002
G_c	Charm quark Yukawa coupling	0.0072 ± 0.0006
G_s	Strange quark Yukawa coupling	0.0006 ± 0.0002
G_t	Top quark Yukawa coupling	1.002 ± 0.029
G_b	Bottom quark Yukawa coupling	0.026 ± 0.003
$\sin \theta_{12}$	Quark CKM matrix angle	0.2243 ± 0.0016
$\sin \theta_{23}$	Quark CKM matrix angle	0.0413 ± 0.0015
$\sin \theta_{13}$	Quark CKM matrix angle	0.0037 ± 0.0005
δ_{13}	Quark CKM matrix phase	1.05 ± 0.24
θ_{qcd}	CP-violating QCD vacuum phase	$< 10^{-9}$
G_{ν_e}	Electron neutrino Yukawa coupling	$< 1.7 \times 10^{-11}$
G_{ν_μ}	Muon neutrino Yukawa coupling	$< 1.1 \times 10^{-6}$
G_{ν_τ}	Tau neutrino Yukawa coupling	< 0.10
$\sin \theta'_{12}$	Neutrino MNS matrix angle	0.55 ± 0.06
$\sin 2\theta'_{23}$	Neutrino MNS matrix angle	≥ 0.94
$\sin \theta'_{13}$	Neutrino MNS matrix angle	≤ 0.22
δ'_{13}	Neutrino MNS matrix phase	?
ρ_Λ	Dark energy density	$(1.25 \pm 0.25) \times 10^{-123}$
ξ_b	Baryon mass per photon ρ_b/n_γ	$(0.50 \pm 0.03) \times 10^{-28}$
ξ_c	Cold dark matter mass per photon ρ_c/n_γ	$(2.5 \pm 0.2) \times 10^{-28}$
ξ_ν	Neutrino mass per photon $\rho_\nu/n_\gamma = \frac{3}{11} \sum m_{\nu_i}$	$< 0.9 \times 10^{-28}$
Q	Scalar fluctuation amplitude δ_H on horizon	$(2.0 \pm 0.2) \times 10^{-5}$
n_s	Scalar spectral index	0.98 ± 0.02

$$C = h = G = k_b = q_e = 1$$

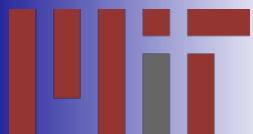
Multiverse level 2: other post-inflationary regions



(Pics from Andrei Linde)

Features:

- Perhaps different *effective* laws of physics (physical constants, particles, symmetries, dimensionality)
- Perhaps even uncountably infinite (Compare a literally parallel Universe; living on another brane)



DO ANY OF THESE QUESTIONS CONFUSE YOU?

1. What is the Universe expanding into?
2. How can stuff be more than 14 billion light years away when the Universe is only 14 billion light years old?
3. Where in space did the Big Bang explosion happen?
4. Did the Big Bang happen at a single point?
5. How could a the Big Bang create an infinite space in a finite time?
6. How could space not be infinite?
7. If the Universe is only 10 billion years old, how can we see objects that are now 30 billion light years away?
8. Don' t galaxies receeding faster than c violate relativity theory?
9. Are galaxies really moving away from us, or is space just expanding?
10. Is the Milky Way expanding?
11. Do we have evidence for a Big Bang singularity?
12. What came before the Big Bang?
13. Should I feel insignificant?

